

Renormalization Group Consistent Treatment of NJL Color-Superconductivity

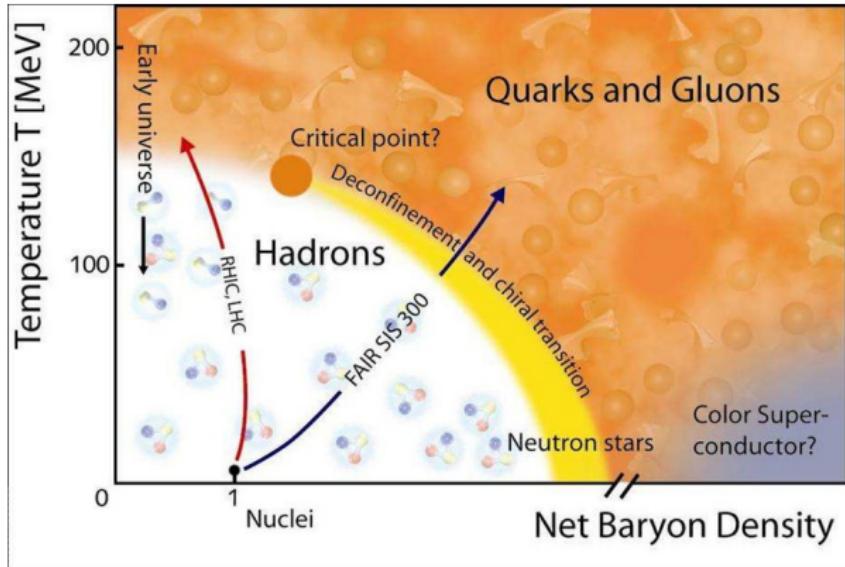
Hosein Gholami in collaboration with M. Hofmann and M. Buballa

Lunch Club Seminar, 8 May 2024, Gießen

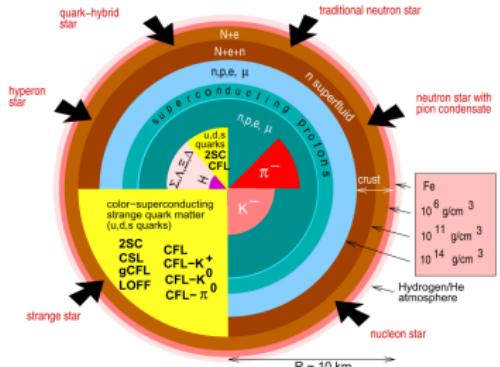


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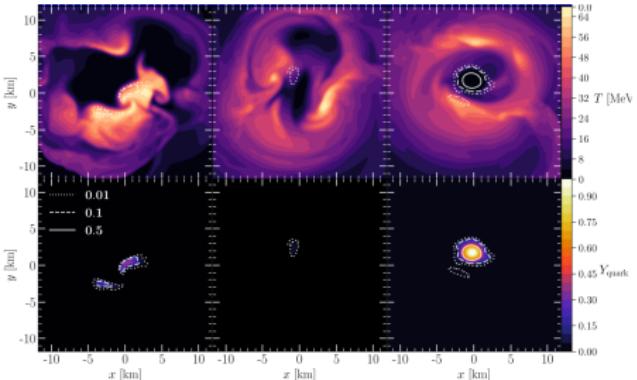




- Phase structure at large (but not asymptotically large) density and moderate T is relevant for neutron stars and neutron star mergers



[Weber (1999)]



[Tootle et al. (2022)]

- ▶ Merger simulations: densities produced in merger remnant might be sufficient to produce quark matter
- ▶ Expected implications: Modified post-merger frequency spectrum [Elias R. Most et al. (2019)] [Bauswein et al. (2019)]

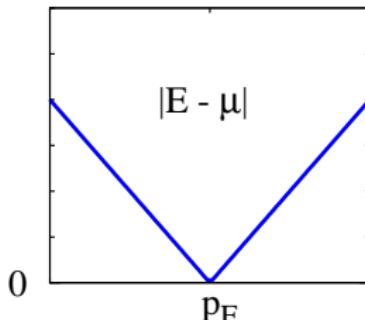
Motivation: signatures of color superconducting phases in neutron star cores or merger remnants?

Method: Use effective models for studying color superconductivity

- ▶ This talk: Treatment of unphysical regularization artefacts in NJL color superconductivity (cut-off artefacts)
- ▶ Investigations of some astrophysical aspects of the model

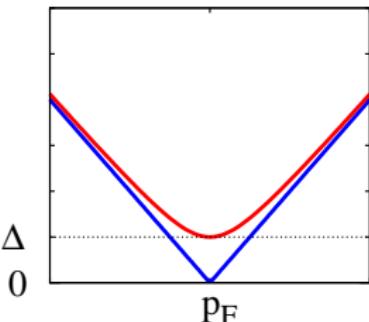
- ▶ Noninteracting fermions at $T = 0$
 - Particles at the Fermi surface can be created at the Fermi surface with no free-energy cost.

- ▶ Cooper theorem: With a finite attractive interaction between particles
 - Fermi surface becomes unstable against pair creation → "Cooper pairs"
 - Bose condensation of the Cooper pairs
⇒ Energy gap Δ in excitation spectrum:
$$\omega = \sqrt{(E - \mu)^2 + \Delta^2}$$

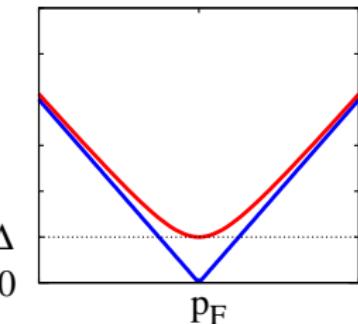


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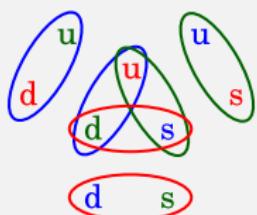
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 \Rightarrow Energy gap Δ in excitation spectrum:
$$\omega = \sqrt{(E - \mu)^2 + \Delta^2}$$
- ▶ **Solid-state superconductor:** Effective attractive interaction between electrons through the interaction with lattice \rightarrow phonons
- ▶ **Color superconductor:** QCD: attractive quark-quark interaction
 - at higher densities via a single gluon exchange in the $\bar{3}$ color channel
 - at intermediate densities also via instanton exchange [Rapp, Schäfer, Shuryak and Velkovsky 1998]



Color Superconductivity

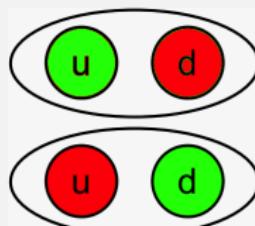
- diquark condensates: $\langle q_i \mathcal{O}_{ij} q_j \rangle$
- Pauli principle: $\mathcal{O} = \mathcal{O}_{\text{spin}} \otimes \mathcal{O}_{\text{color}} \otimes \mathcal{O}_{\text{flavor}} = \text{totally antisymmetric}$
- most attractive channel: color $\bar{3}$ and spin 0 (antisymmetric) \Rightarrow mixing flavors

Color-flavor-locking (CFL)



Large $\mu \gg M_s$
 $SU(3)_{c+L+R} \otimes Z_2$
 3 finite gap parameters
 $\Delta_{ud}, \Delta_{us}, \Delta_{ds}$

SC



Intermediate $\mu \lesssim M_s$
 $SU(2)_c \otimes SU(2)_L \otimes SU(2)_R \otimes$
 $U(1)_{\bar{B}} \otimes Z_2$
 1 finite gap parameter Δ_{ud}

Nambu Jona-Lasinio (NJL)-type model

$$\mathcal{L} =$$

$$\bar{\psi}(i\cancel{\partial} - \textcolor{blue}{m})\psi \quad \text{kinetic term}$$

$$+ \textcolor{blue}{G} \sum \left[(\bar{\psi}\tau_a\psi)^2 + (\bar{\psi}i\gamma_5\tau_a\psi)^2 \right] \quad \text{scalar NJL interaction}$$

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$$- \textcolor{blue}{K} [\det_f(\bar{\psi}(\mathbb{1} + \gamma_5)\psi) + \det_f(\bar{\psi}(\mathbb{1} - \gamma_5)\psi)] \quad \text{'t Hooft (KMT) interaction}$$

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$$- \textcolor{blue}{K} [\det_f(\bar{\psi}(1 + \gamma_5)\psi) + \det_f(\bar{\psi}(1 - \gamma_5)\psi)] \quad \text{'t Hooft (KMT) interaction}$$

$$+ G \eta_D \sum (\bar{\psi}i\gamma_5\tau_A\lambda_{A'}\psi^c)(\bar{\psi}^c i\gamma_5\tau_A\lambda_{A'}\psi) \quad \text{diquark interaction}$$

with charge conjugated spinor $\psi^c = C\bar{\psi}^T$

Nambu Jona-Lasinio (NJL)-type model

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 & - \textcolor{blue}{K} [\det_f(\bar{\psi}(\mathbb{1} + \gamma_5)\psi) + \det_f(\bar{\psi}(\mathbb{1} - \gamma_5)\psi)] && \text{'t Hooft (KMT) interaction} \\
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 \end{aligned}$$

with charge conjugated spinor $\psi^c = C\bar{\psi}^T$

Mean field approximation: Linearise theory around condensates

$$\phi_f = \langle \bar{\psi}_f \psi_f \rangle \quad f = u, d, s$$

$$\Delta_A = -2G\eta_D \langle \bar{\psi}^c \gamma_5 \tau_A \lambda_A \psi \rangle \quad A = 2(ud), 5(us), 7(ds)$$

Then minimizing with respect to these condensates

- Λ' , G , K , m fitted to vacuum meson spectrum
- Regularization: sharp 3-momentum cutoff $\Lambda' = 602 \text{ MeV}$

- ▶ Chemical potential matrix in color-flavor space:

$$\mu_{f,c} = \mu + Q_f \mu_Q + \lambda_{3,c} \mu_3 + \lambda_{8,c} \mu_8$$

$$\text{e.g. } \mu_{u,r} = \mu + \frac{2}{3} \mu_Q + \mu_3 + \frac{1}{\sqrt{3}} \mu_8$$

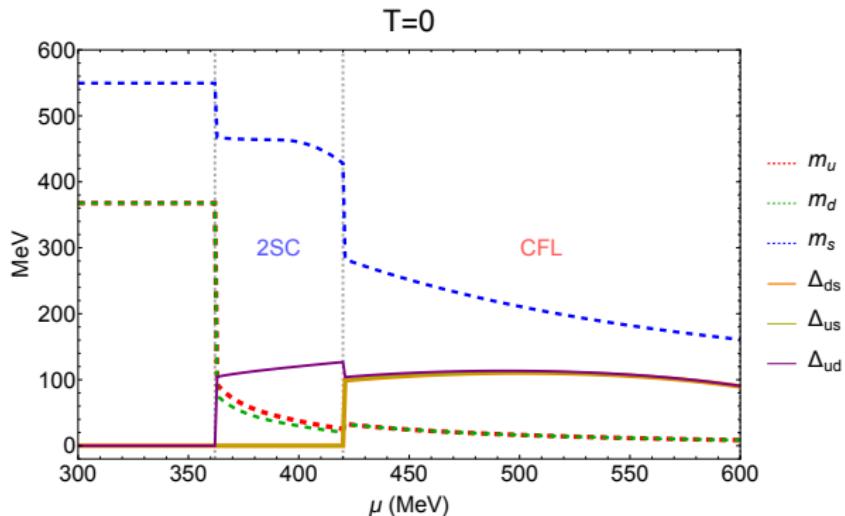
- ▶ Neutron star: Enforce charge and color-neutrality locally, i.e. for every phase:

$$\frac{\partial \Omega}{\partial \mu_Q} = \frac{\partial \Omega}{\partial \mu_3} = \frac{\partial \Omega}{\partial \mu_8} = 0$$

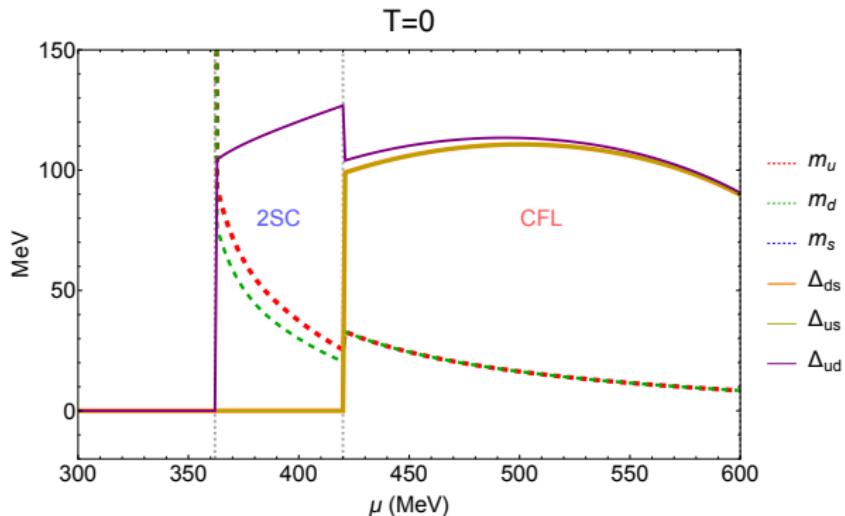
- ▶ Leptonic contribution: e^- and μ^- in β -equilibrium $\mu_e = \mu_\mu = -\mu_Q$
- ▶ Optimization problem with nonlinear constraints
 - 3 quark mass gap equations
 - 3 diquark gap equations
 - 3 neutrality constraints
 - in total 9 equations to be solved self-consistently

Condensates in Mean Field

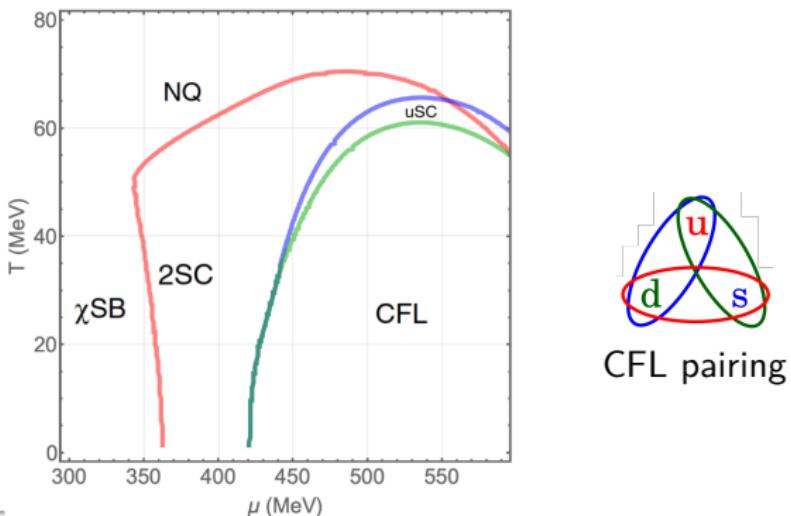
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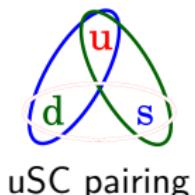
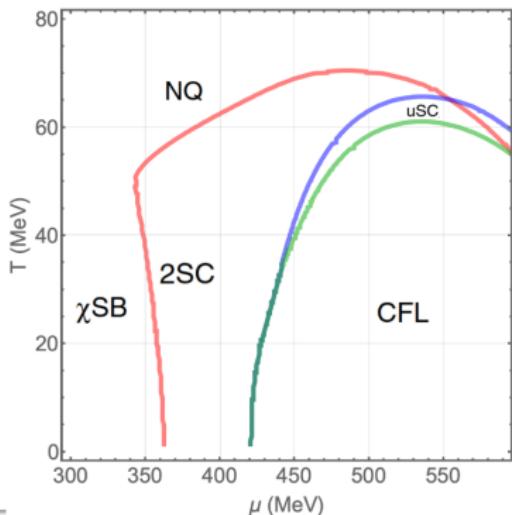


- ▶ Cutoff artefact: Δ values descend at high chemical potentials/number densities



- ▶ Cutoff artefacts
 - Gaps and phase boundary to normal phase decrease in value at $\mu \sim \Lambda'$
 - Appearance of uSC phase [Fukushima 2005]
- ▶ Previous explanations for uSC: $T=0$ arguments → Not relevant for $T \neq 0$

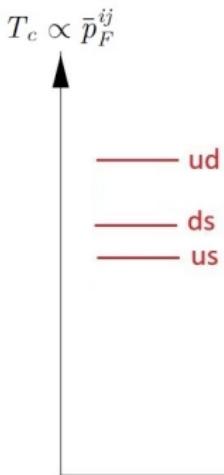
Phase Diagram of Neutral Quark Matter



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► **Puzzle:** Absence of expected dSC phase in CFL melting pattern [Iida et al 2004]

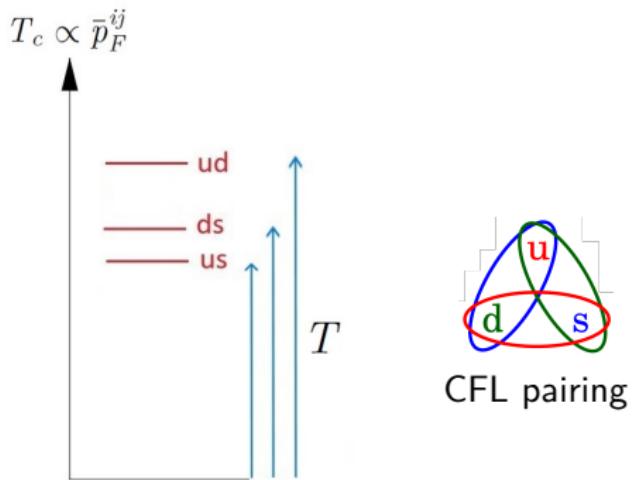
- Ginzburg-Landau analysis around T_c
- T_c of a pairing is proportional to average Fermi momenta of that pairs:
$$T_c \propto \bar{p}_F^{ij} = \frac{1}{2}(p_F^i + p_F^j)$$



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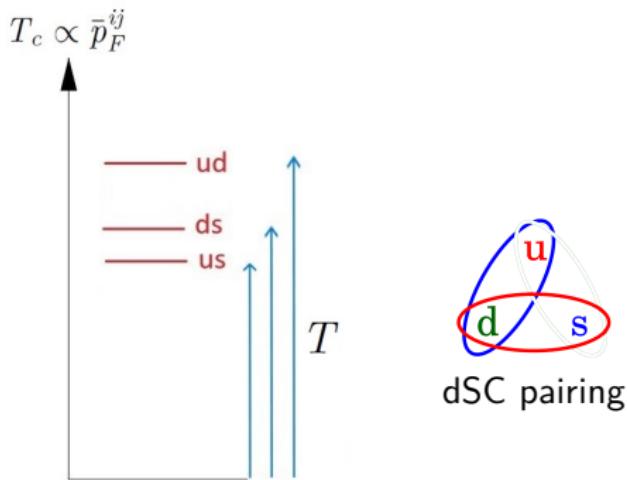


- Melting pattern: CFL → dSC → 2SC

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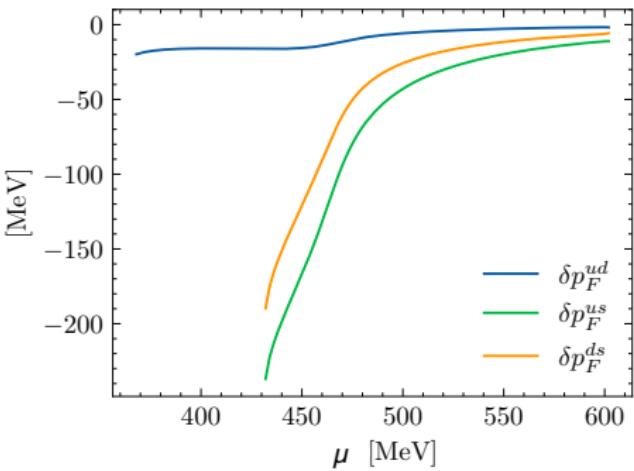
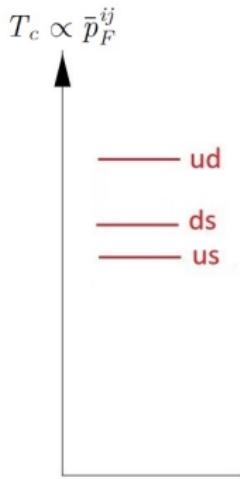


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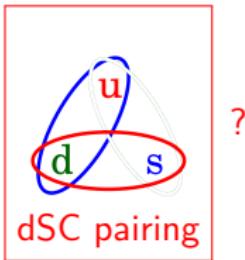
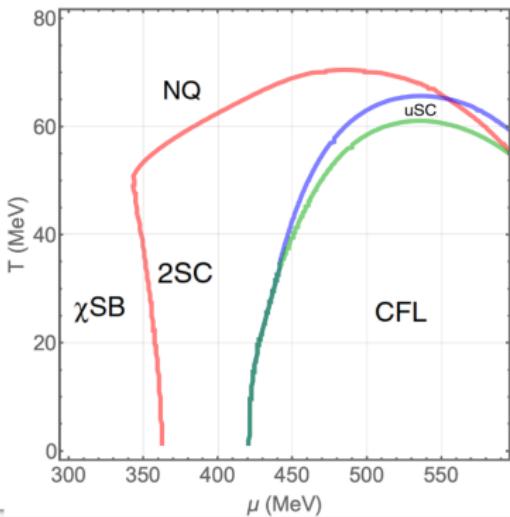
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$$T_c \propto \bar{p}_F^{ij} = \frac{1}{2}(p_F^i + p_F^j) = \mu + \delta p_F^{ij}$$



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Solution to Regularization artefacts:

- ▶ Use *renormalization group-consistent* regularization presented by Braun et al (2016)

SciPost

SciPost Phys. 6, 056 (2019)

Renormalization group consistency and low-energy effective theories

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Technische Universität Darmstadt, D-64289 Darmstadt, Germany

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3 Institut für Theoretische Physik, Universität Heidelberg,
Philosophenweg 16, D-69120 Heidelberg, Germany

- ▶ Scale dependent effective action Γ_k includes all physics above momentum scale k
- ▶ *Full Quantum effective action* Γ includes all scales $\Gamma := \lim_{k \rightarrow 0} \Gamma_k$
- ▶ Renormalization group calculation: Ansatz for theory at scale Λ is evolved to $k = 0$ by solving the RG-flow equation

Renormalization Group Consistency

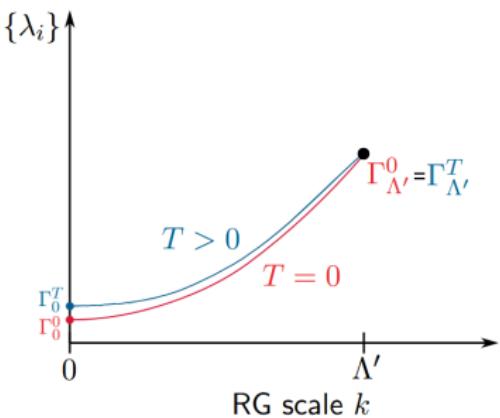
The full quantum effective action must not depend on the initial scale $k = \Lambda$

$$\Lambda \frac{d\Gamma}{d\Lambda} = 0.$$

Solution of the flow equation in mean field

$$\Gamma(\mu, T) = \Gamma_\Lambda(\mu, T) + \frac{1}{2} \int_{p < \Lambda} \frac{d^3 p}{(2\pi)^3} T \cdot \text{Tr} \log \frac{S^{-1}}{T} \quad (1)$$

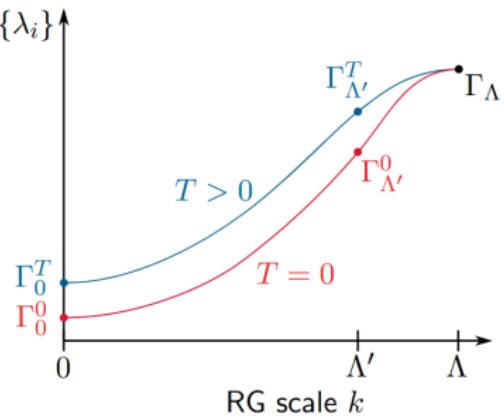
Problem: Λ' isn't big enough with respect to all scales of the system when in medium: $\Lambda' \sim \mu, T, \Delta_i, M_j$.



Solution of the flow equation in mean field

$$\Gamma(\mu, T) = \Gamma_\Lambda(\mu, T) + \frac{1}{2} \int_{p < \Lambda} \frac{d^3 p}{(2\pi)^3} T \cdot \text{Tr} \log \frac{S^{-1}}{T} \quad (2)$$

Assumption: Λ bigger than all scales of the system: $\Lambda \gg \mu, T, \Delta_i, M_j, \Lambda'$. Then $\Gamma_\Lambda(\mu, T) \approx \Gamma_\Lambda(\mu = 0, T = 0)$ can be calculated from flow equation in vacuum



- ▶ In most of the mean field cases: RG consistent treatment is equivalent to only regularizing the vacuum part

$$\Gamma(\mu, T, \chi) = \frac{1}{2\pi^2} \int_0^{\Lambda'} dp p^2 \sum_j \left(\epsilon_j(\mu, \chi) + 2T \ln \left(1 + e^{-\frac{\epsilon_j(\mu, \chi)}{T}} \right) \right) - \mathcal{V}$$

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- ▶ If $\epsilon_j(\mu, \chi) = \tilde{\epsilon}_j(\chi) \pm \mu$

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- ▶ In most of the mean field cases: RG consistent treatment is equivalent to only regularizing the vacuum part → separation of vacuum and medium
- ▶ In the case of color superconductivity: vacuum part can't be separated from the medium part

$$\epsilon_{\pm} = \sqrt{(E \pm \mu)^2 + \Delta^2}$$

$$\Gamma(\mu, T, \chi) \supset \frac{1}{2\pi^2} \int_0^\Lambda dp p^2 (\epsilon_+ + \epsilon_-) \supset \mu^2 \Delta^2 \log \Lambda$$

- ▶ In most of the mean field cases: RG consistent treatment is equivalent to only regularizing the vacuum part → separation of vacuum and medium
- ▶ In the case of color superconductivity: vacuum part can't be separated from the medium part → issue of medium divergence
- ▶ 3-flavour NJL color-superconductivity suffers from a medium divergence

$$\mu^2(\Delta_{ud}^2 + \Delta_{us}^2 + \Delta_{ds}^2) \log \Lambda$$

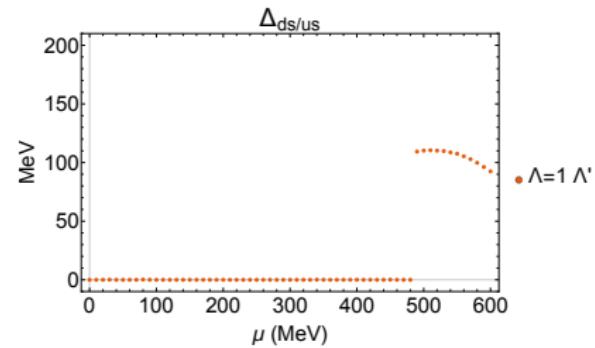
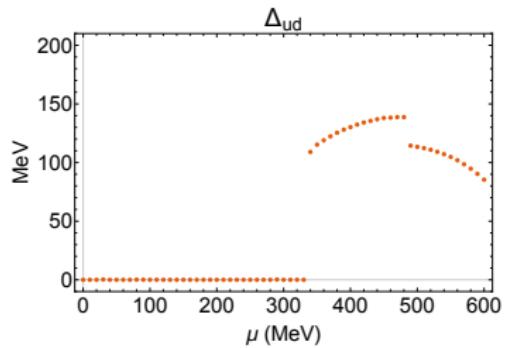
this divergence can be balanced by including a counterterm

$$Y_k(\Delta)\mu^2 \quad \text{with} \quad Y_k(\Delta) = 0 \text{ for } k < \Lambda'$$

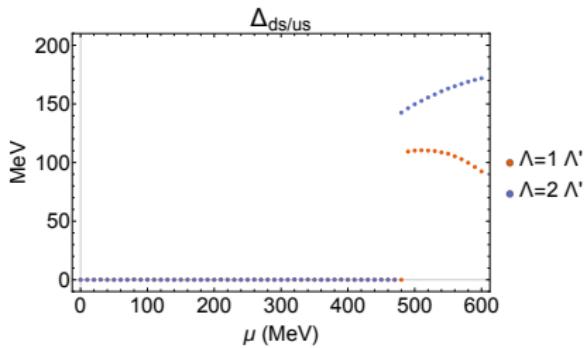
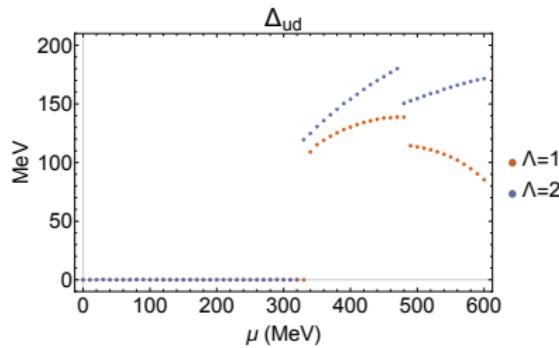
in the Lagrangian: renormalization of the medium part → Note that we are not renormalizing a QFT!

- ▶ This choice of counter-term is equivalent to regularizing the corresponding term raised in the expansion of the medium part

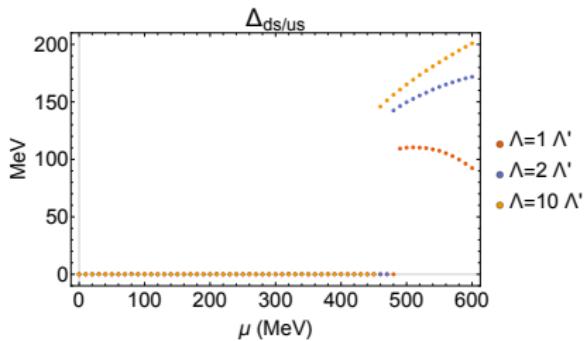
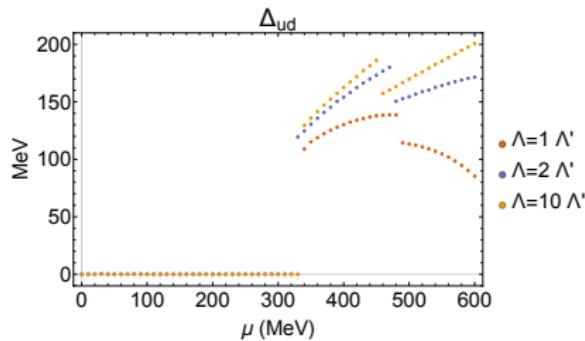
Flowing up to the original cut-off scale $\Lambda = \Lambda'$ (i.e. not RG consistent)



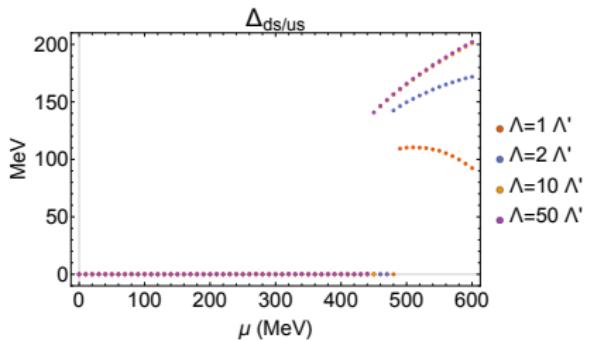
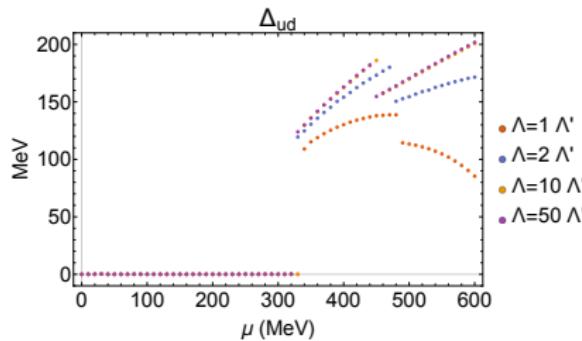
Flowing up to $\Lambda = 2\Lambda'$



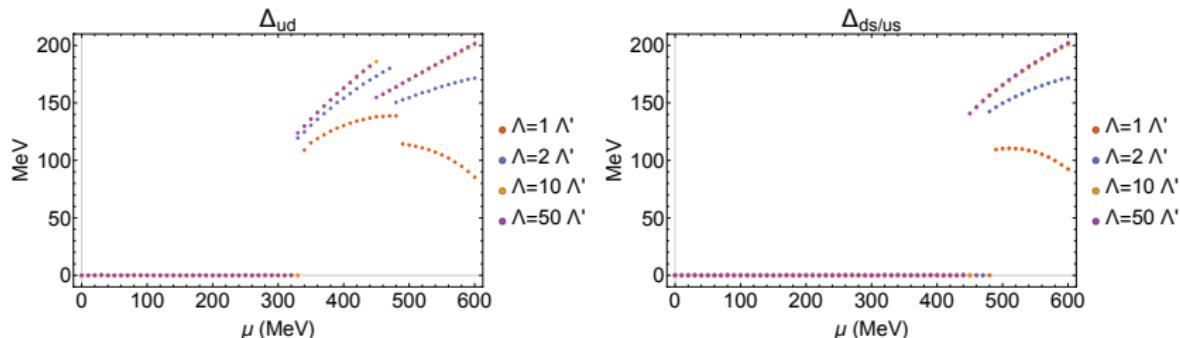
Flowing up to $\Lambda = 10\Lambda'$



Flowing up to $\Lambda = 50\Lambda'$



Condensates in RG-Consistent Model



- ▶ At $\Lambda \approx 10\Lambda'$, results become almost independent of Λ
- ▶ Expected Increasing trend \rightarrow cut-off artefacts removed
- ▶ Gap values become enlarged for the same diquark coupling

- ▶ For a generalized chemical potential analytically derived form of divergence

$$\left(\Delta_{ud}^2 \bar{\mu}_{ud}^2 + \Delta_{us}^2 \bar{\mu}_{us}^2 + \Delta_{ds}^2 \bar{\mu}_{ds}^2 \right) \log \Lambda$$

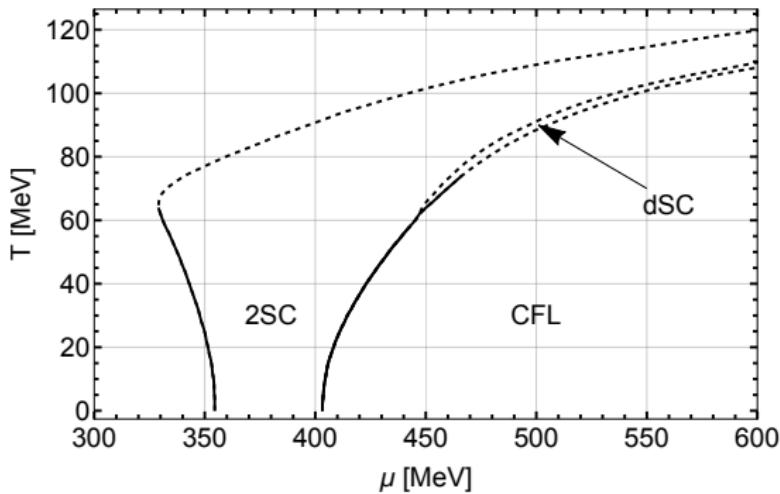
with $\bar{\mu}_{ij}$ being the average chemical potential of quark flavour species i and j

- ▶ These divergences should be removed for an RG consistent model
- ▶ Behaviour of divergence suggests that

$$\mathcal{L}^R \supset \frac{1}{2} Y_{ud}(\Delta_{ud}) \bar{\mu}_{ud}^2 + \frac{1}{2} Y_{ds}(\Delta_{ds}) \bar{\mu}_{ds}^2 + \frac{1}{2} Y_{us}(\Delta_{us}) \bar{\mu}_{us}^2$$

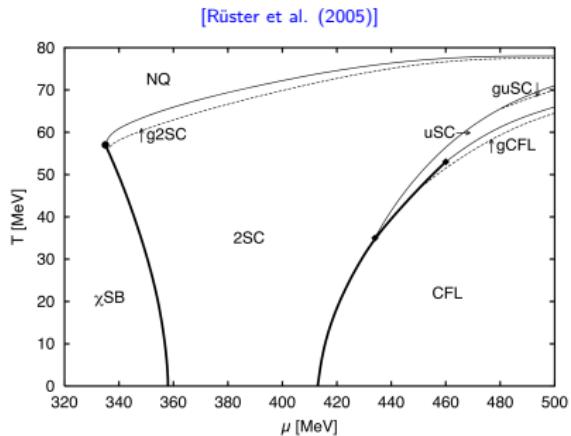
- ▶ One can analytically drive the expression for these 3 renormalization factors

Results: Phase Diagram

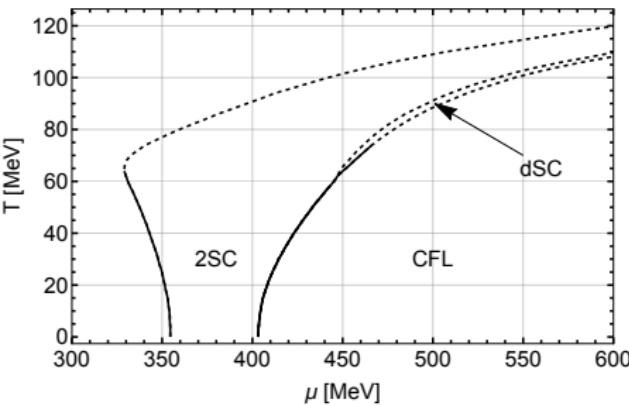


- ▶ Cut-off artefacts are removed
 - Expected increasing trend of phase boundaries
 - No uSC phase
- ▶ Expected dSC phase appears in CFL melting: **Puzzle solved**

Results: Phase Diagram



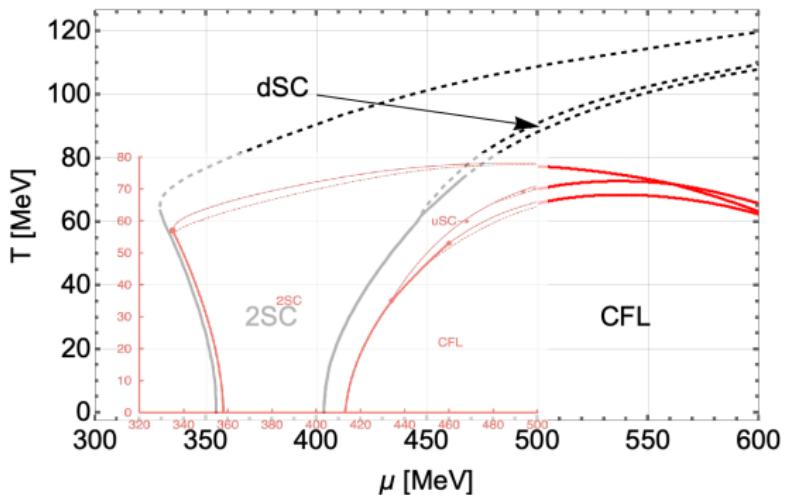
not RG consistent



RG consistent $\Lambda = 10\Lambda'$

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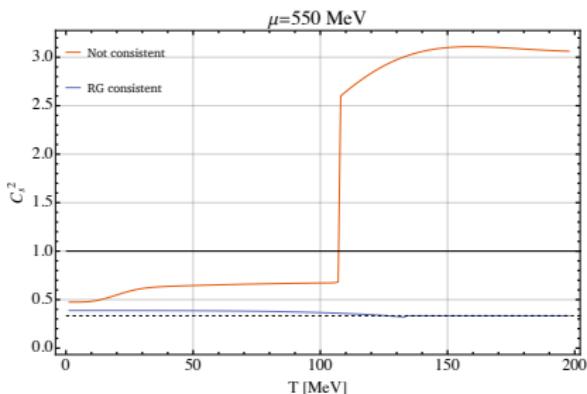
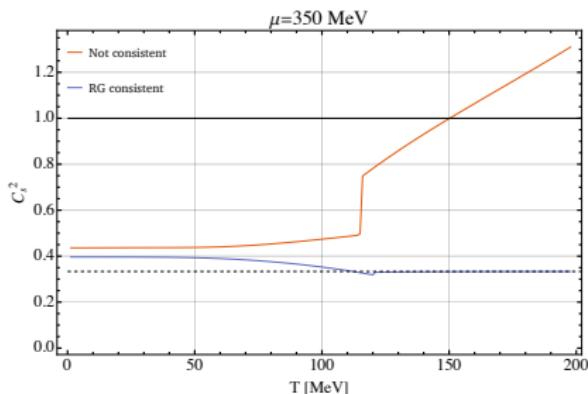
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 - Expected increasing trend of phase boundaries
 - No uSC phase
- ▶ Expected dSC phase appears in CFL melting: **Puzzle solved**

Thermodynamics

- ▶ Pressure: $P(T, \mu) = \frac{T}{V_3} \Gamma(T, \mu)$
- ▶ Equation of State: Energy density $\epsilon = -P + \mu n + sT$ as a function of P
- ▶ Connects microphysics to astrophysics of neutron stars via TOV equation → Mass-radius curves
- ▶ **Example:** Speed of sound indicates the stiffness of the EOS: Higher C_s means stiffer EOS → higher TOV mass
 - $C_s^2 = \frac{dP}{d\epsilon} \Big|_s$
 - Causality: $C_s^2 < 1$
 - Chiral limit: $C_s^2 = 1/3$

A natural outcome of the RG consistent treatment: correct thermodynamic limits

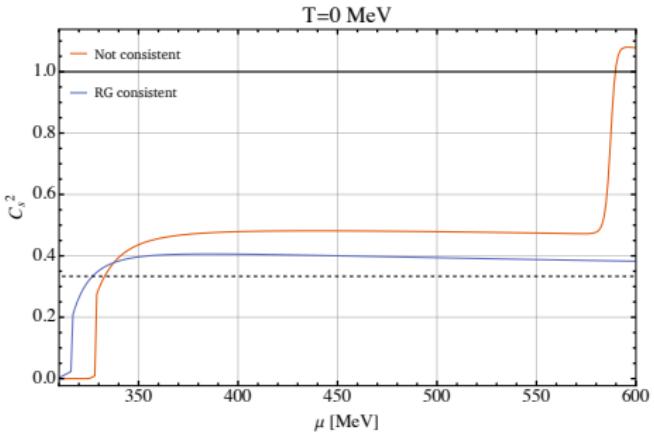
- ▶ Correct thermodynamics are mostly important for the astrophysics related calculations
- ▶ Above the T_c at high densities speed of sound should go to the chiral limit $C_s^2 = 1/3$
 - can only be achieved if the medium part is not regularized
 - RG consistent treatment automatically satisfies this



Results: Thermodynamics

A natural outcome of the RG consistent treatment: correct thermodynamic limits

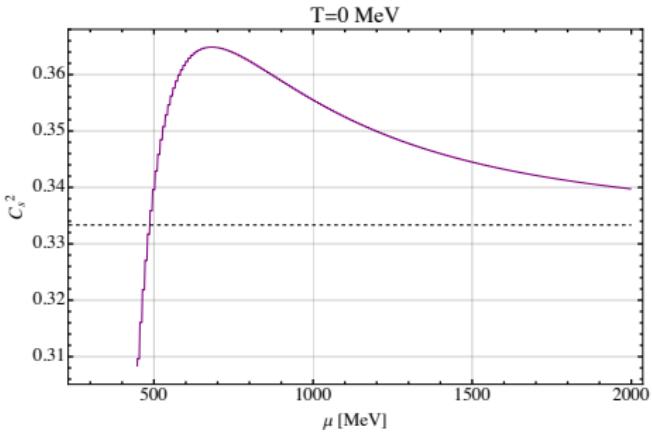
- ▶ Correct thermodynamics are mostly important for the astrophysics related calculations
- ▶ One expects at higher densities: $P \approx \alpha\mu^4 + \beta\mu^2\Delta^2$
 - if $\Delta < \mu \Rightarrow C_s^2 \rightarrow 1/3$ at high densities



Results: Thermodynamics

A natural outcome of the RG consistent treatment: correct thermodynamic limits

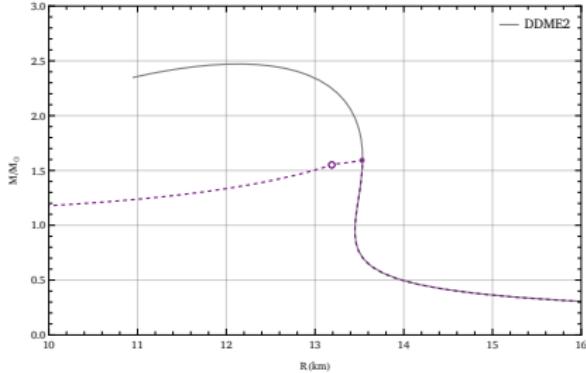
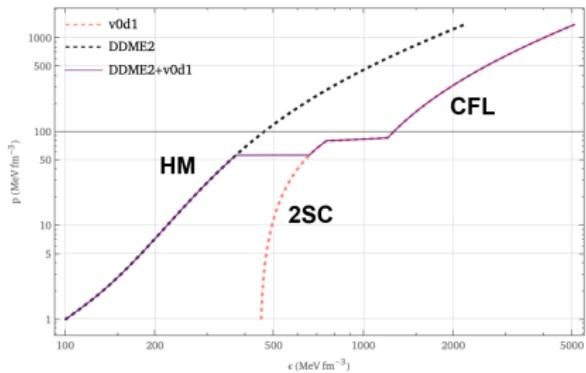
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Results: MR curves

Procedure:

- ▶ Choose EoS for Hadronic Matter (HM) at low densities that satisfies all observational constraints (here: relativistic mean-field model with density-dependent couplings DDME2 [Lalazissis et al. (2005)])
- ▶ Calculate Maxwell construction: $\{P, \mu_B, T\}_{\text{HM}} = \{P, \mu_B, T\}_{\text{QM}}$ in β -equilibrium at the point of the phase transition
By construction, this gives a first order phase transition from HM to QM
- ▶ Calculate M-R-relation



Model: Vector interaction

Vector interaction provides stiffness of the equation of state at high temperatures to reach $2M_{\odot}$ hybrid stars [Klähn et al (2007, 2013), Alaverdyan (2022)]

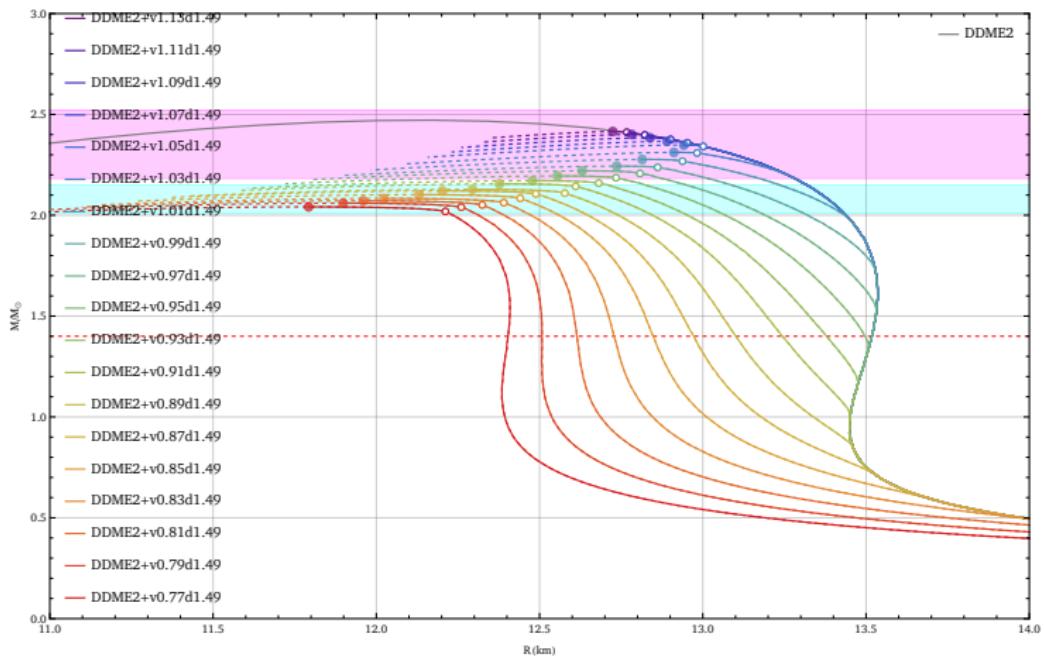
$$\mathcal{L} =$$

$$\begin{aligned}
 & \bar{\psi}(i\not{\partial} - \textcolor{blue}{m})\psi && \text{kinetic term} \\
 & + \textcolor{blue}{G} \sum \left[(\bar{\psi}\tau_a\psi)^2 + (\bar{\psi}i\gamma_5\tau_a\psi)^2 \right] && \text{scalar NJL interaction} \\
 & - \textcolor{blue}{K} [\det_f(\bar{\psi}(1 + \gamma_5)\psi) + \det_f(\bar{\psi}(1 - \gamma_5)\psi)] && \text{'t Hooft (KMT) interaction} \\
 & + G \eta_D \sum (\bar{\psi}i\gamma_5\tau_A\lambda_{A'}\psi^c)(\bar{\psi}^c i\gamma_5\tau_A\lambda_{A'}\psi) && \text{diquark interaction} \\
 & - G \eta_V (\bar{\psi}\gamma^\mu\psi)^2 && \text{vector interaction}
 \end{aligned}$$

- ▶ 2 free parameters η_D, η_V can be constrained by observational constraints on the static EoS of isolated hybrid stars

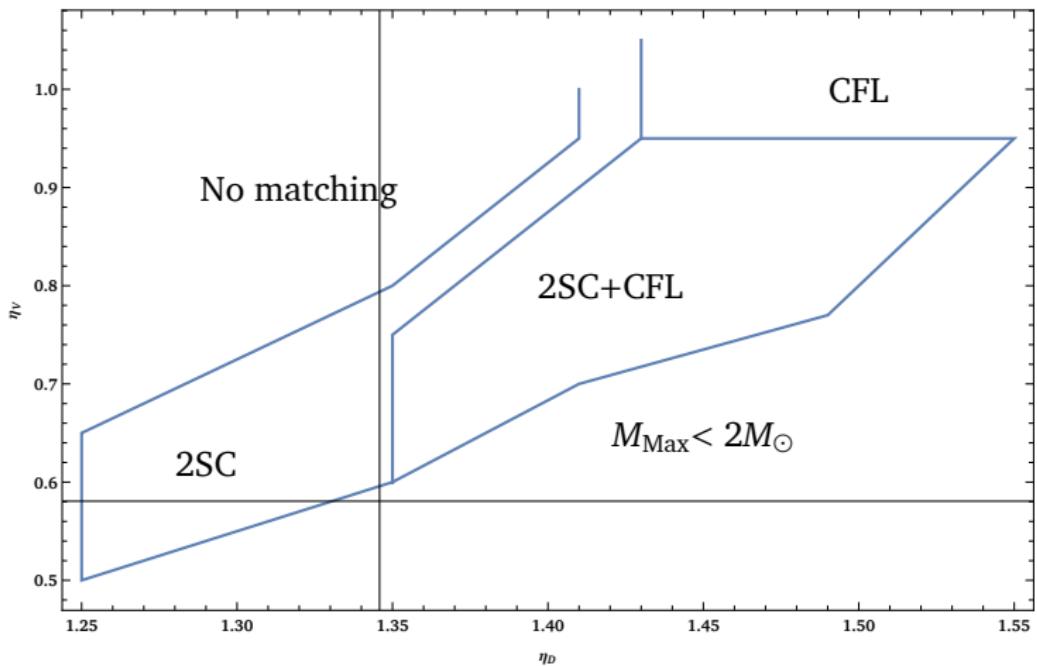
Results: MR curves

Variation of the vector coupling at constant $\eta_D = 1.49$



Message: Model allows for hybrid EoS containing stable 2SC, 2SC+CFL and CFL-cores consistent with the observation of $2M_\odot$ neutron stars

Work in preparation



Summary

- ▶ NJL color-superconductivity suffers from cut-off artefacts
- ▶ RG-consistent formulation systematically removes the cutoff artefacts and changes the phase diagram in terms of critical temperatures, diquark condensate values and phase transition points
- ▶ RG-consistent formulation for neutral CSC matter is in agreement with expected dSC phase in CFL melting pattern
- ▶ Different type of possible phase transitions to CSC matter in agreement with astrophysical observations
- ▶ We can limit free parameters of the model by the astrophysical constraints

Outlook

- ▶ Main interest: Study imprints of color superconductivity in neutron star mergers
- ▶ Publishing results...

Thank you for listening!