

Field experience with SiPMs in Cherenkov and Scintillator Detectors

Thomas Bretz



Physics
Institute III

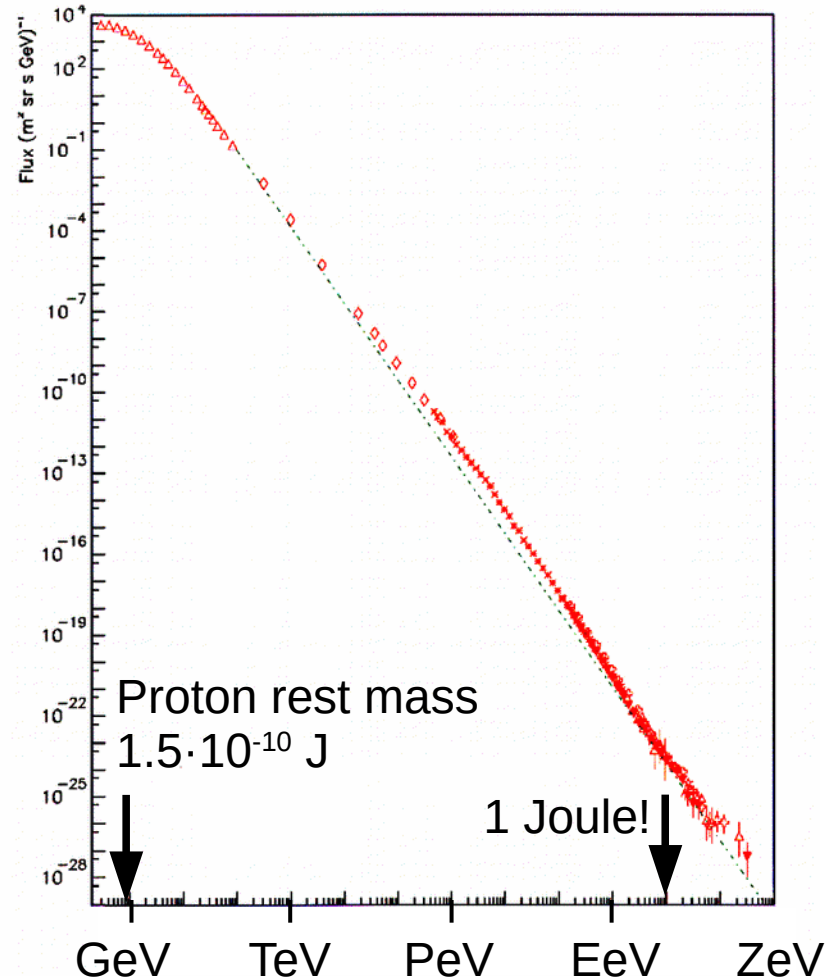
RWTHAACHEN
UNIVERSITY



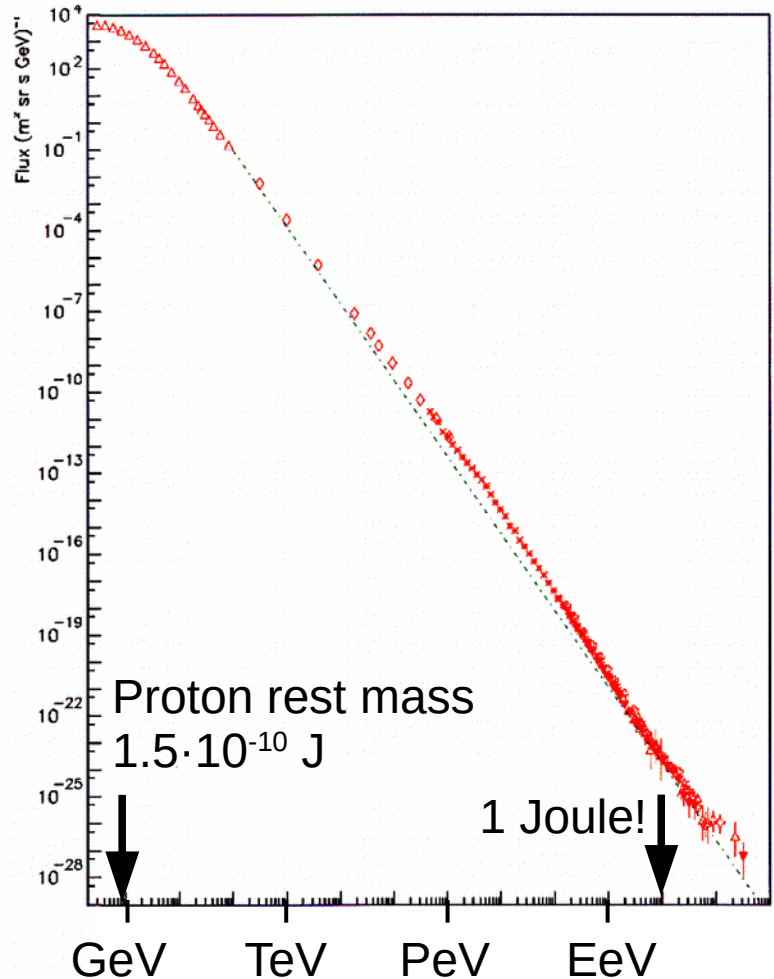
SiPM (A selection)

- High Photo Detection efficiency (up to 50%)
- Very robust (semi-conductor, illumination)
- Single Photon resolution
- **Small sensitive area**
 - significant increase possible by light guides
- **Gain stability**
 - needs voltage regulation
- **Dynamic range limited by number of G-APD cells**
 - needs good understanding of properties for unfolding

The cosmic ray spectrum (all particles)



The cosmic ray spectrum (all particles)



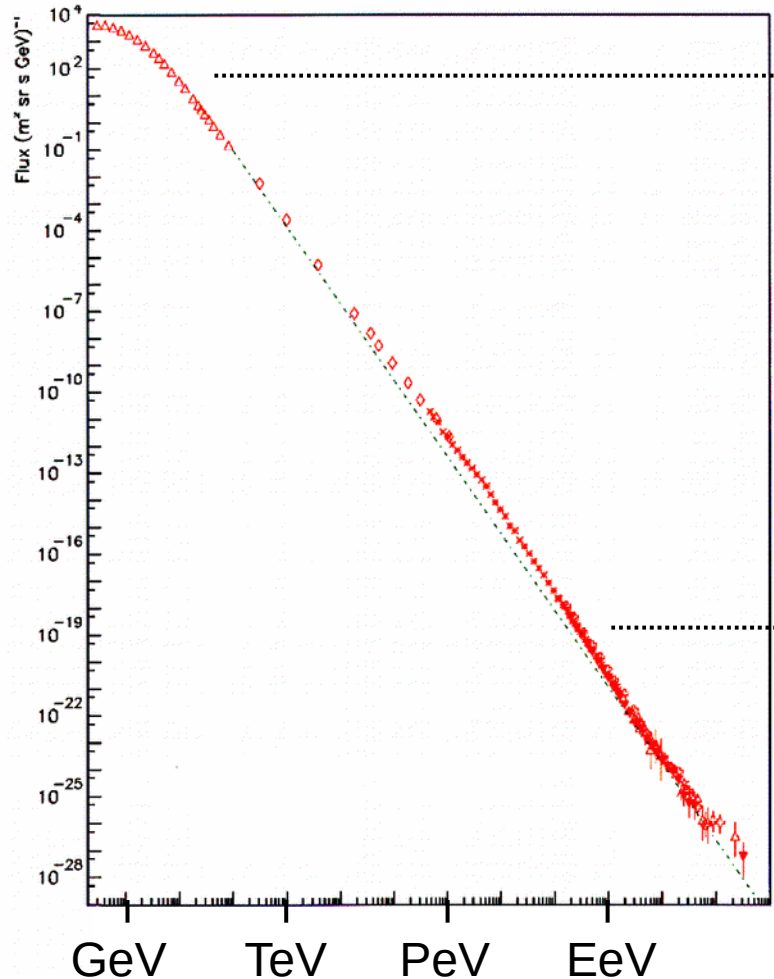
← $F(E > 10^9 \text{ eV}) \sim 1 \text{ particle / s m}^2$
Amazon river: 206 000 000 litre / s



← $F(E > 10^{19.5} \text{ eV}) \sim 1 \text{ particle / century km}^2$
1 drop / month
(1 drop $\sim 5\text{mg}$)



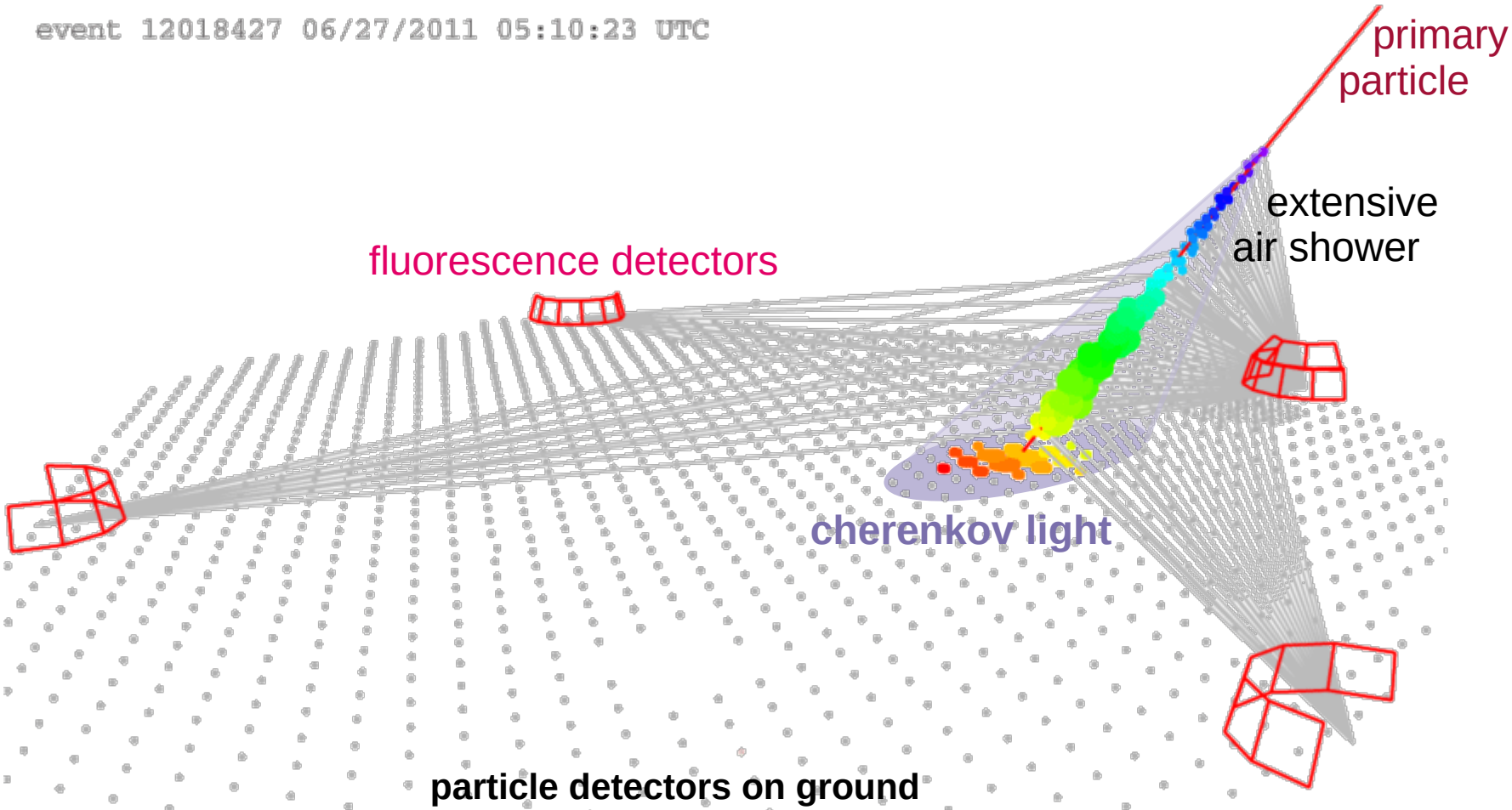
The cosmic ray spectrum



can only
be measured
from ground
through
extensive air
showers

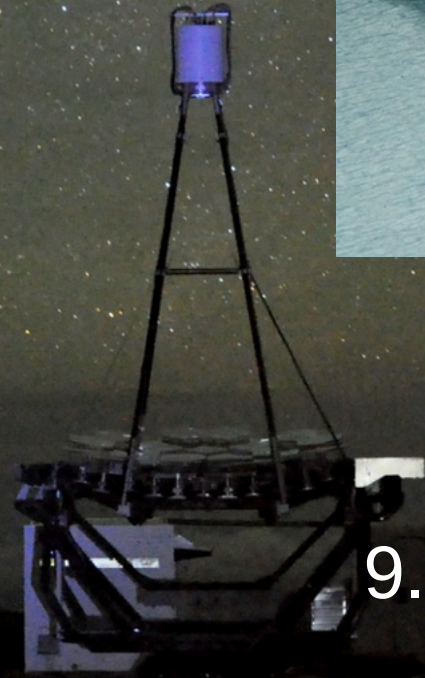
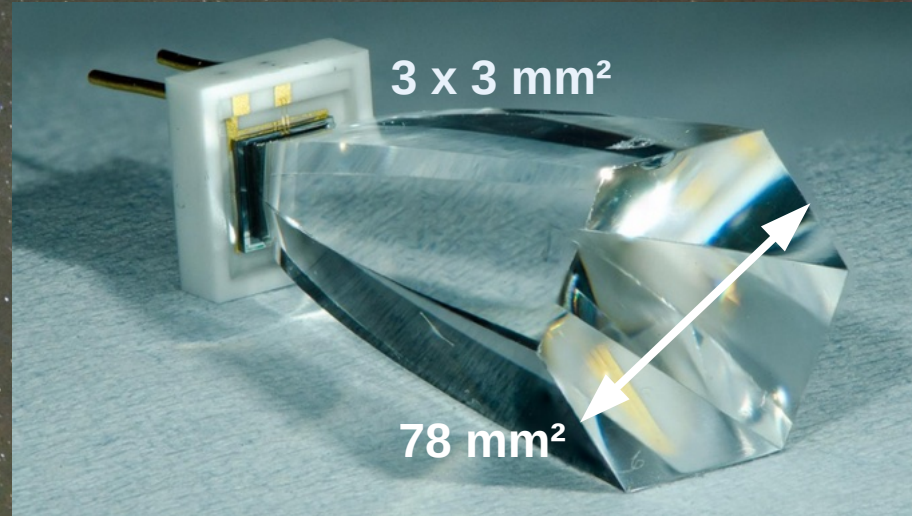
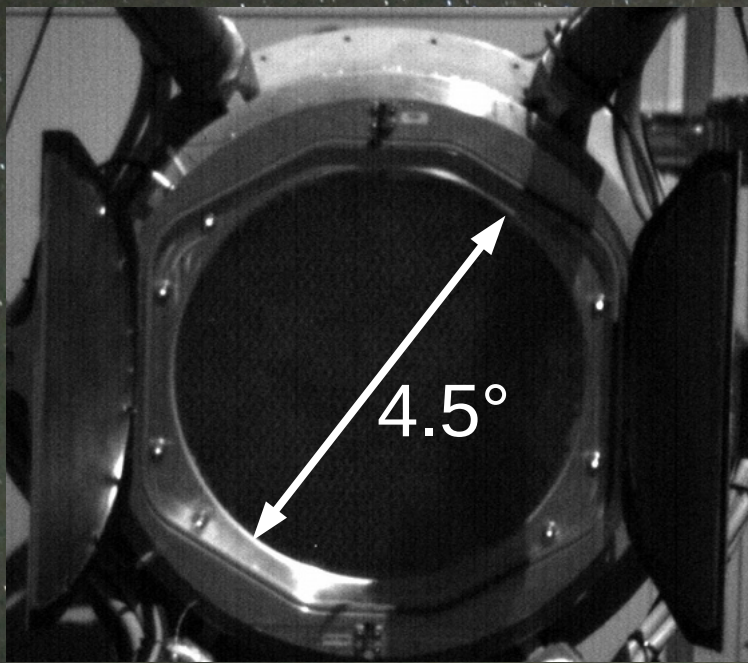
Hunt for the origin of cosmic rays

event 12018427 06/27/2011 05:10:23 UTC



FACT – The First G-APD Cherenkov Telescope

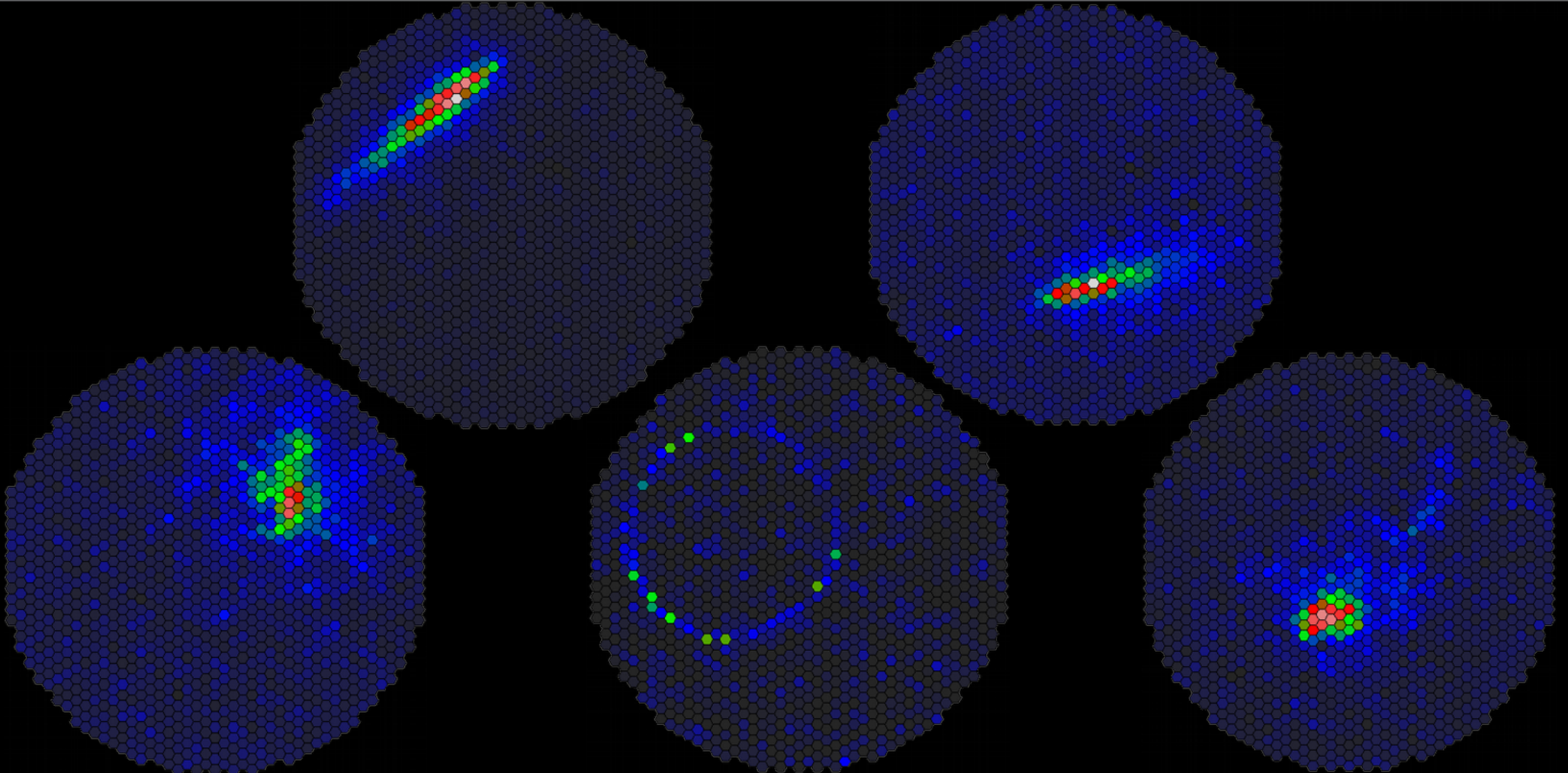
1440 SiPM with solid cone



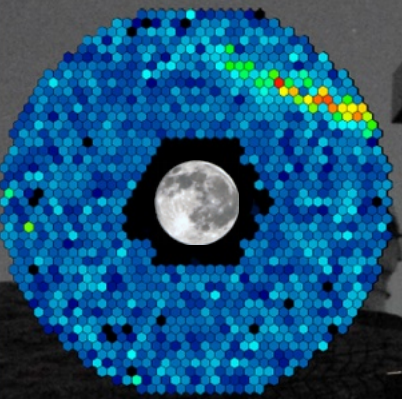
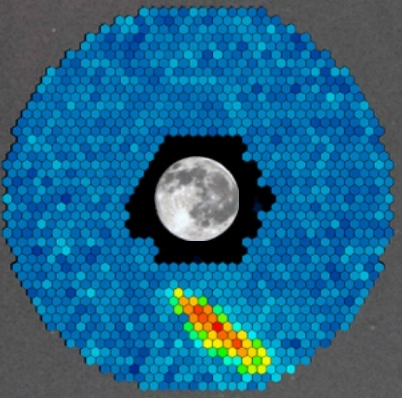
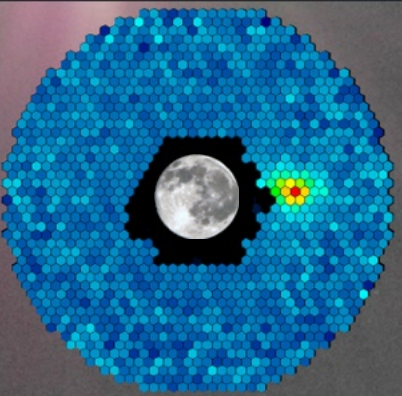
9.5 m²

JINST (2013) 8 P06008
JINST (2014) 9 P10012

Selected events (first light: Oct 11th 2011)

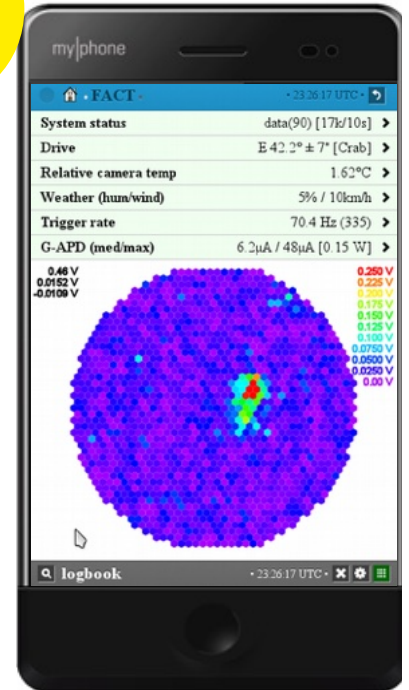


Operation during moon light

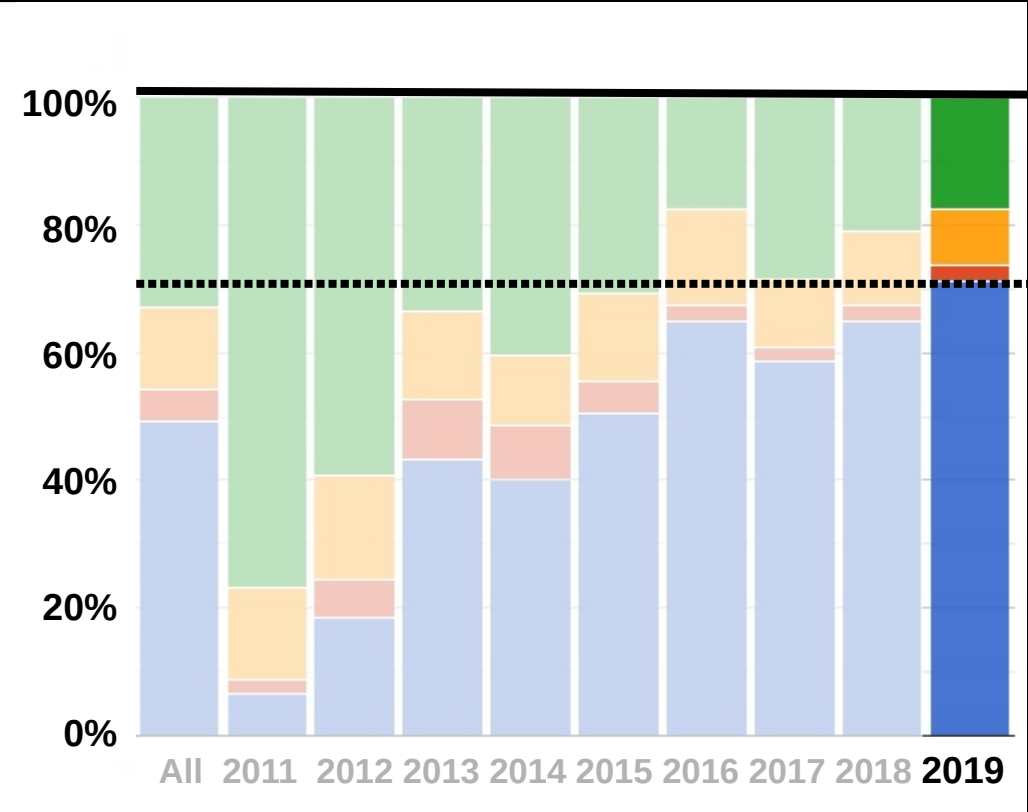


no calibration
required

unattended and
automatic
data taking



www.fact-project.org/smartfact



42% Night time per year

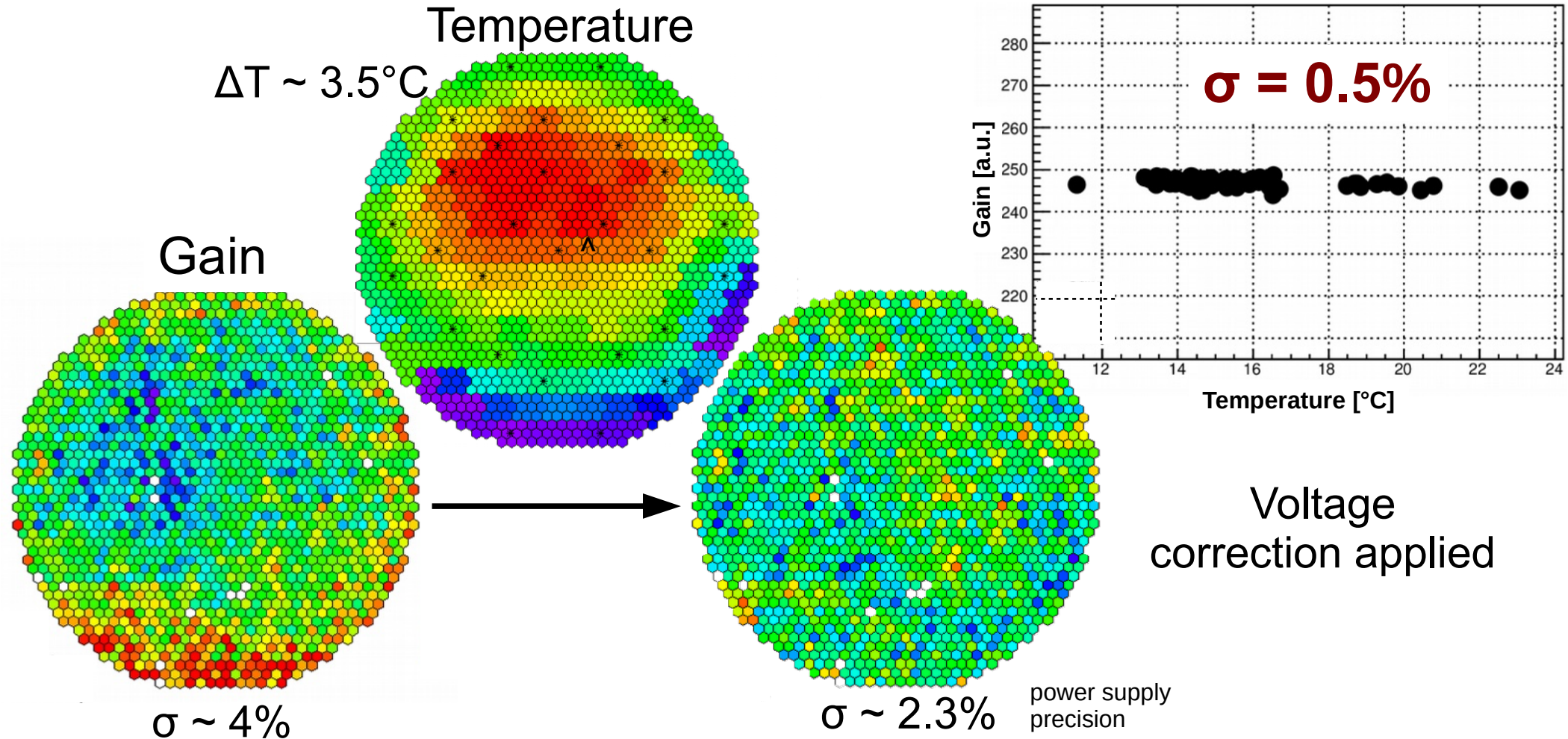
~71% Physics data per year

→ $0.42 \cdot 0.71 = 30\%$
2600 h

possible
→ $x 1.16 = 35\%$
3000 h

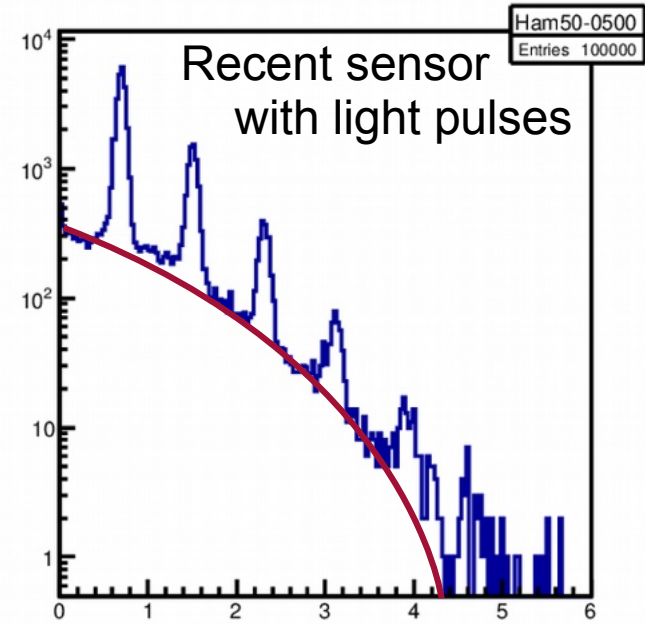
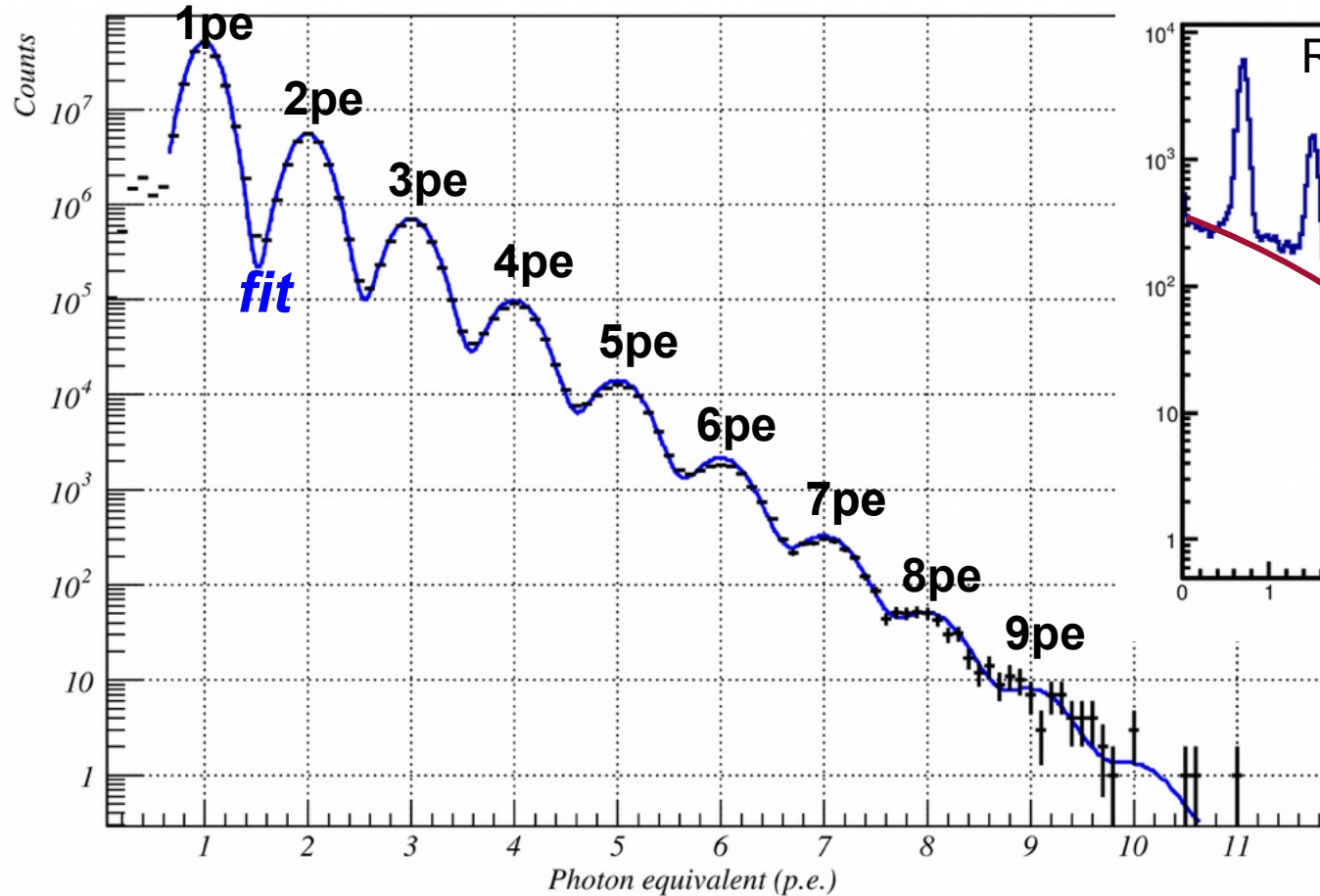
→ ideally suited for monitoring

Temperature and gain distribution in the camera



Self calibrating / Stability

all pixels; one year; temp: $\sim 0^{\circ}\text{C} - 25^{\circ}\text{C}$



A compact and light-weight telescope

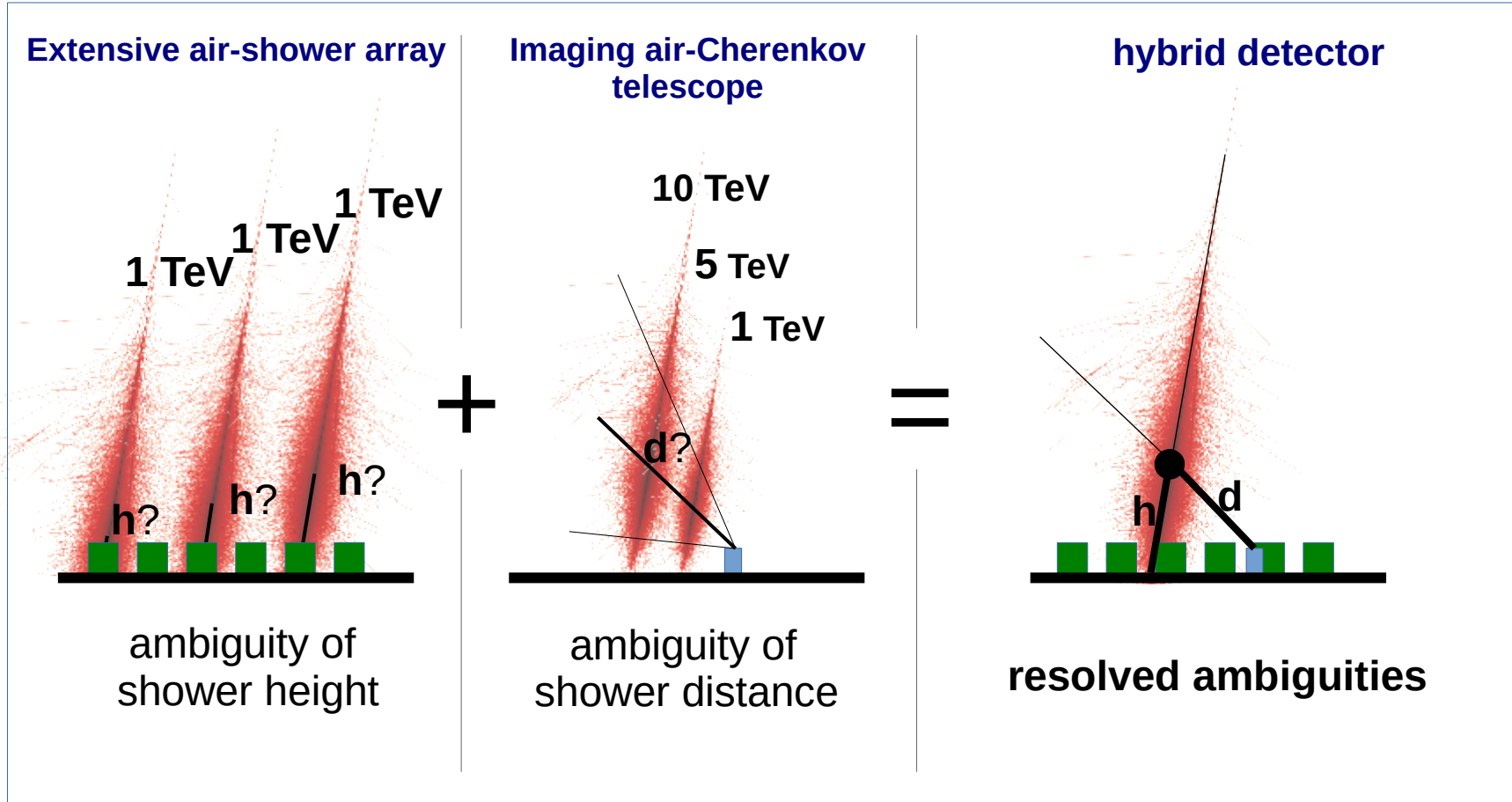
- *A compact and light-weight refractive telescope for the observation of extensive air showers* [JINST \(2018\) 13 P07024](#)



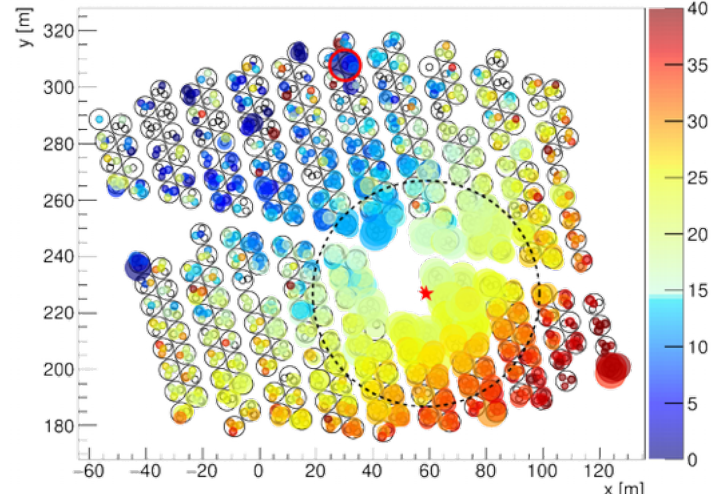
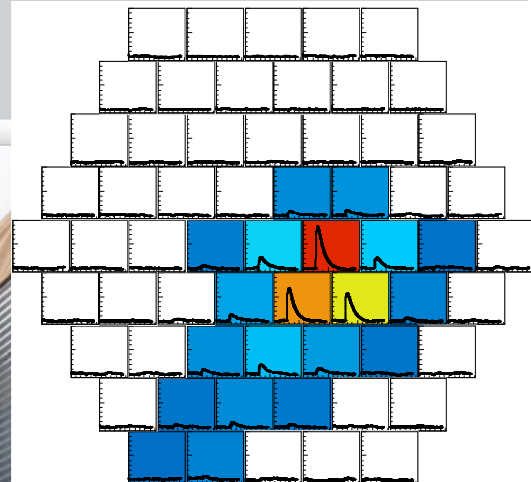
50cm Fresnel lens



The (over-simplified) Cherenkov concept

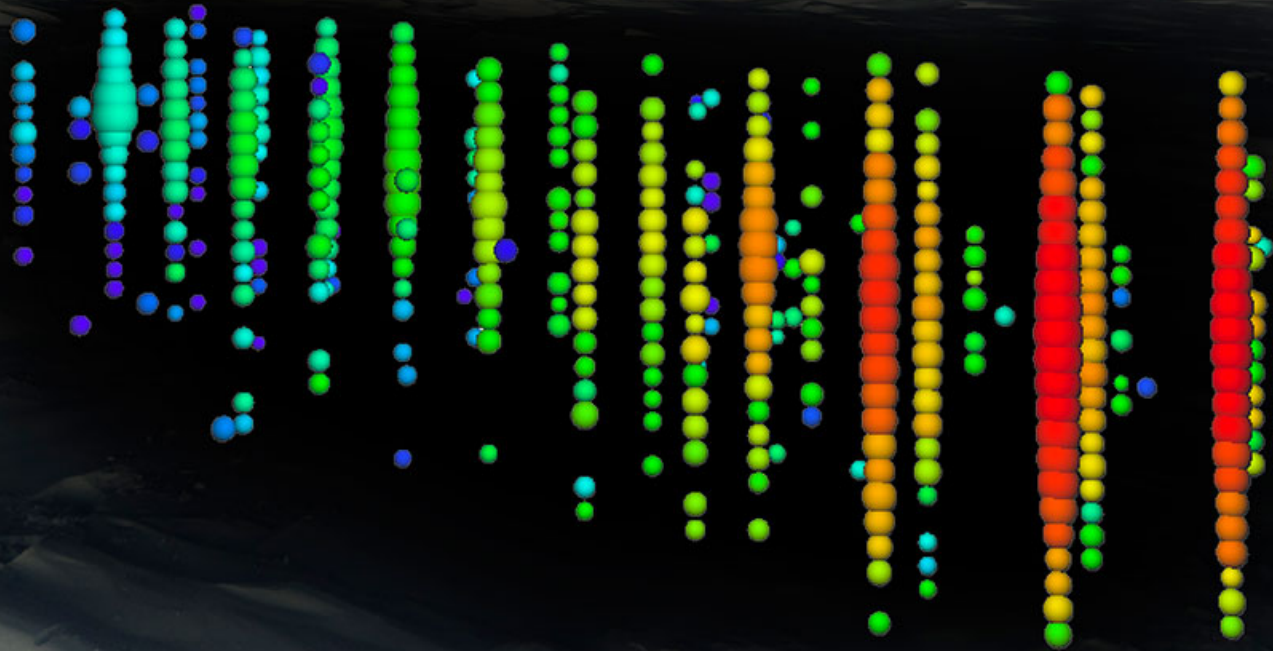


HAWC's Eye – First light

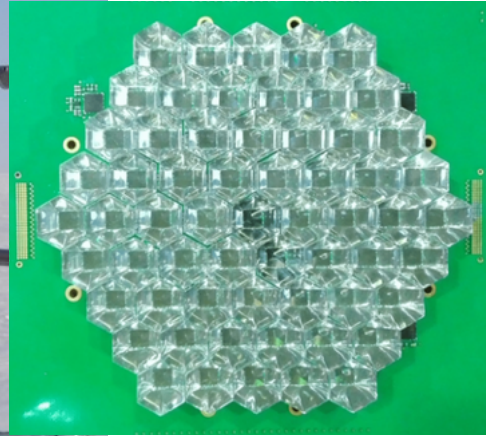
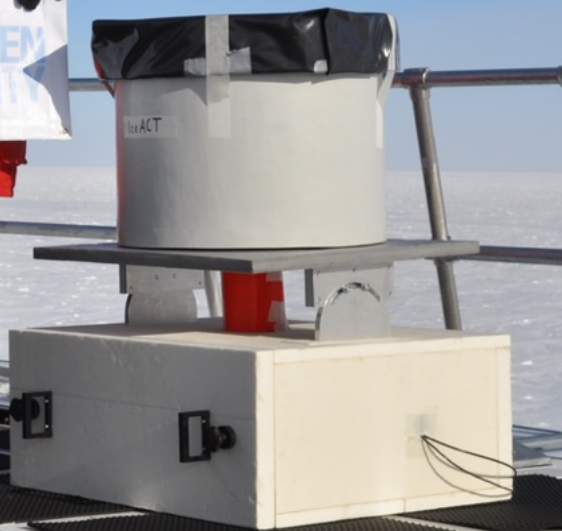
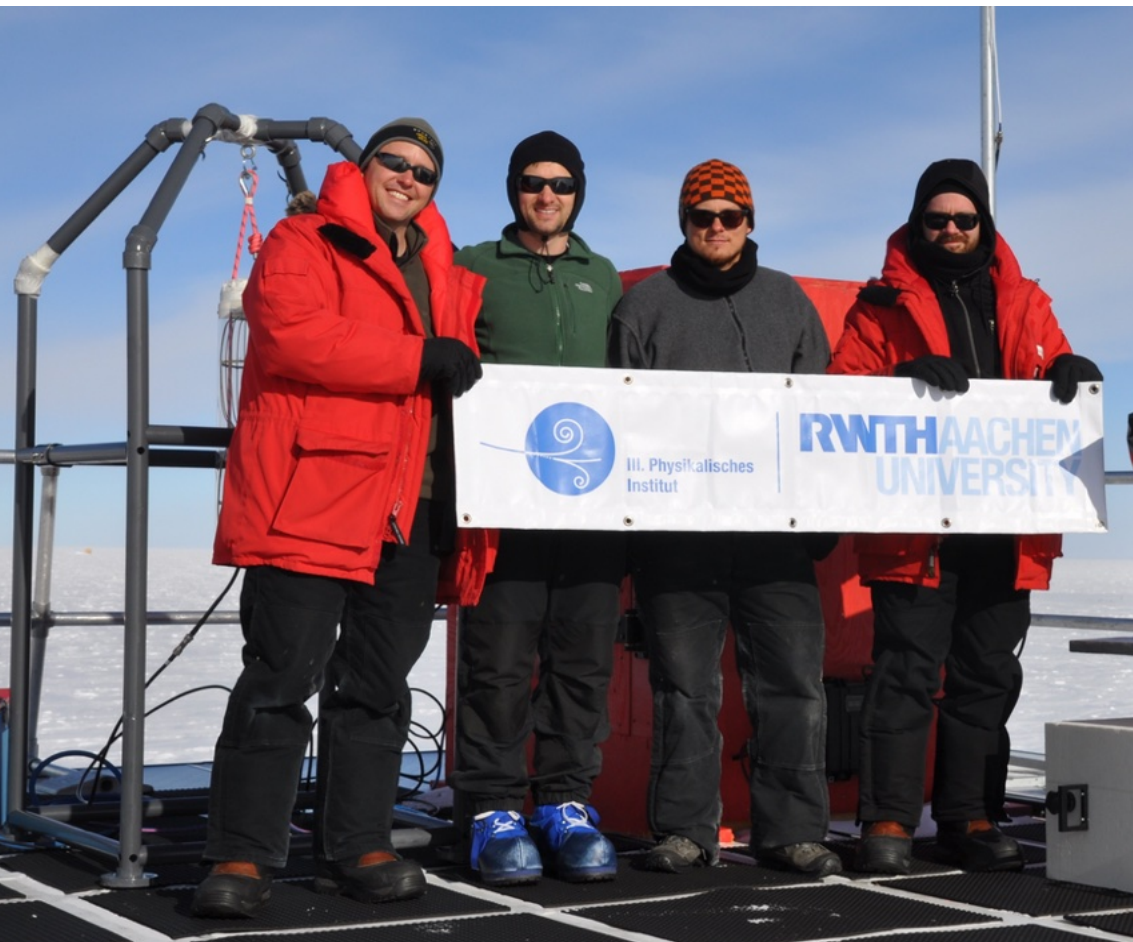


HAWC, Mexico

IceCube @ South Pole



Veto / CR Composition at the knee



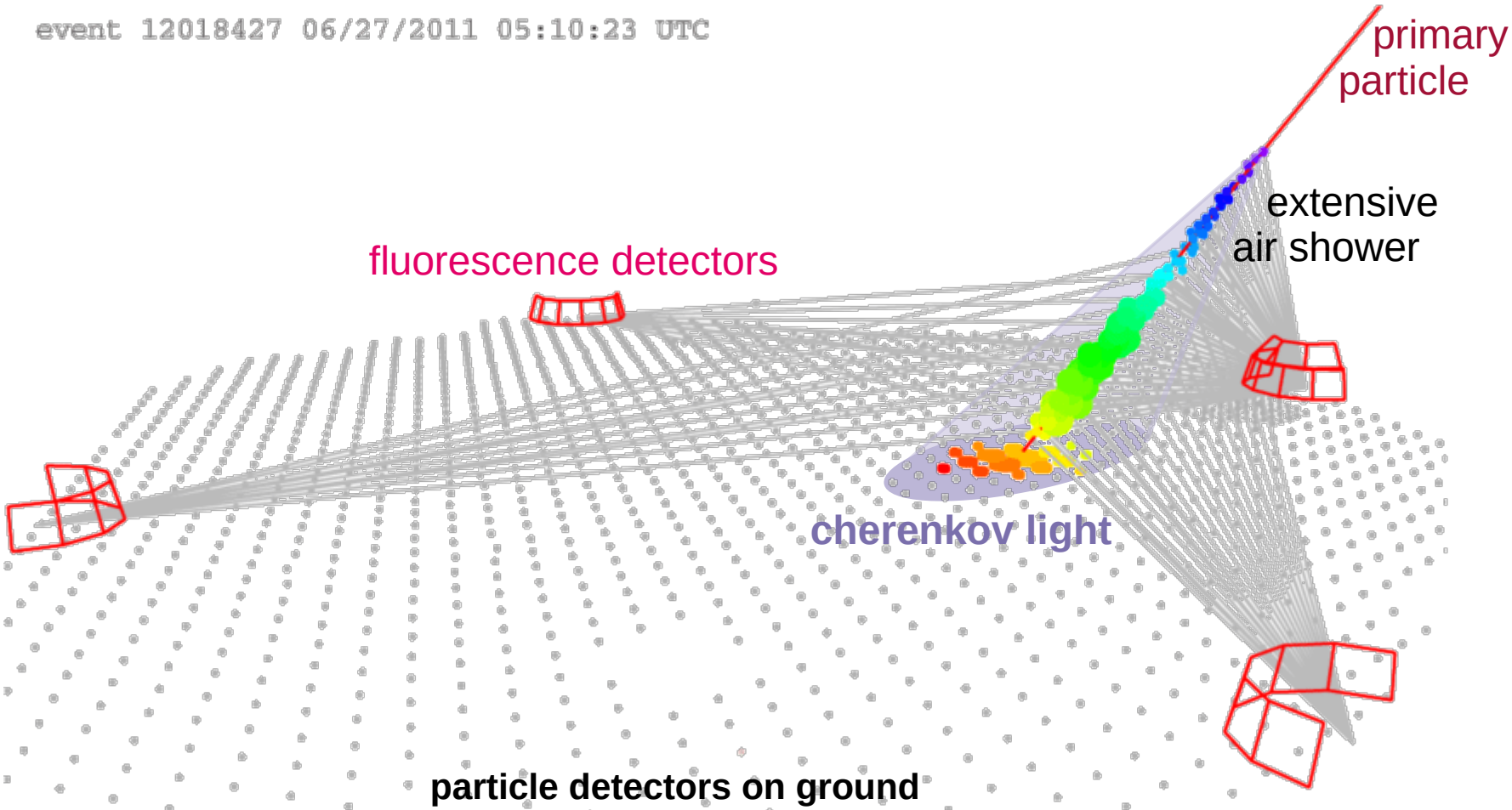
Student lab course astroparticle physics

cosmic ray physics
air shower physics



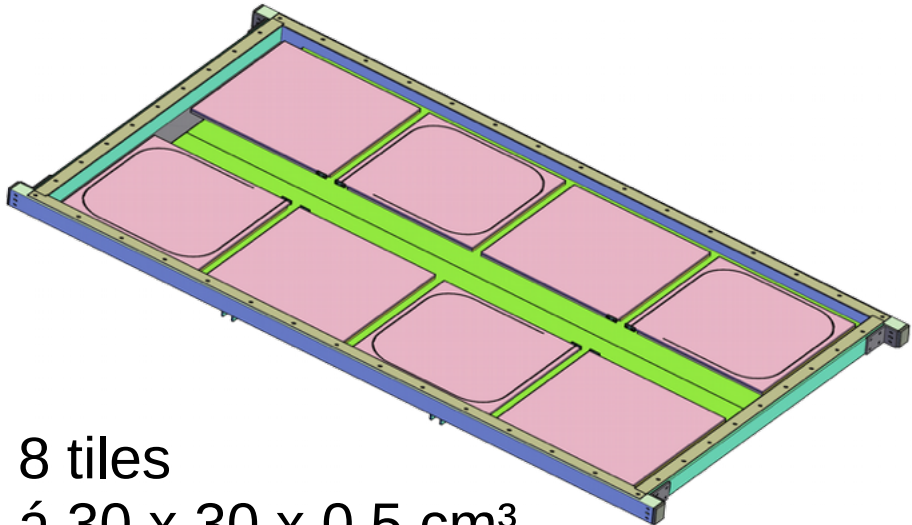
Hunt for the origin of cosmic rays

event 12018427 06/27/2011 05:10:23 UTC

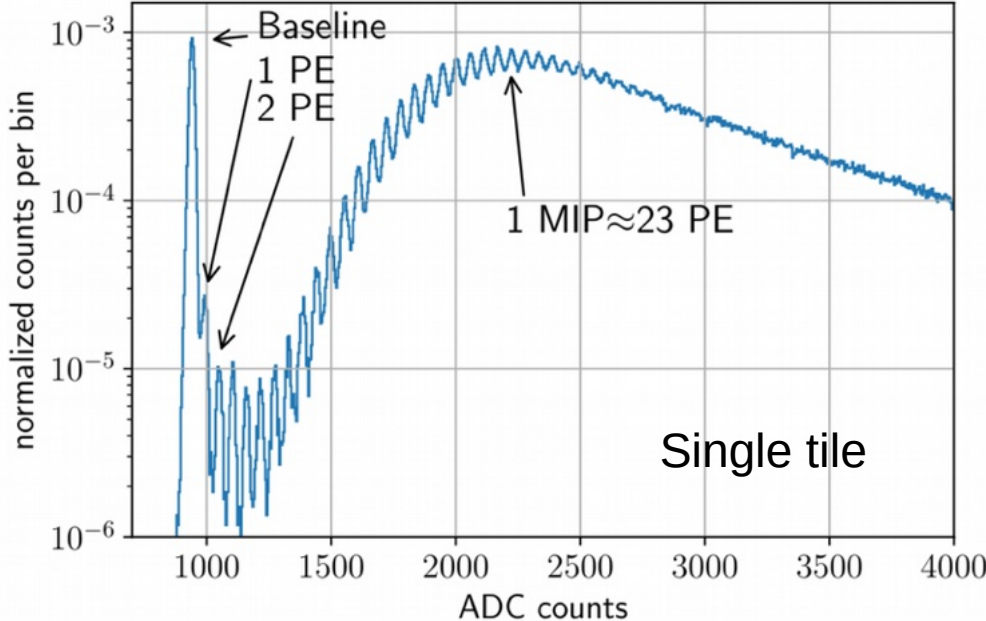


Aachen Muon Detector

- Dedicated application of SiPM
- Wavelength shifting fibers
- One 1mm² SiPM per tile



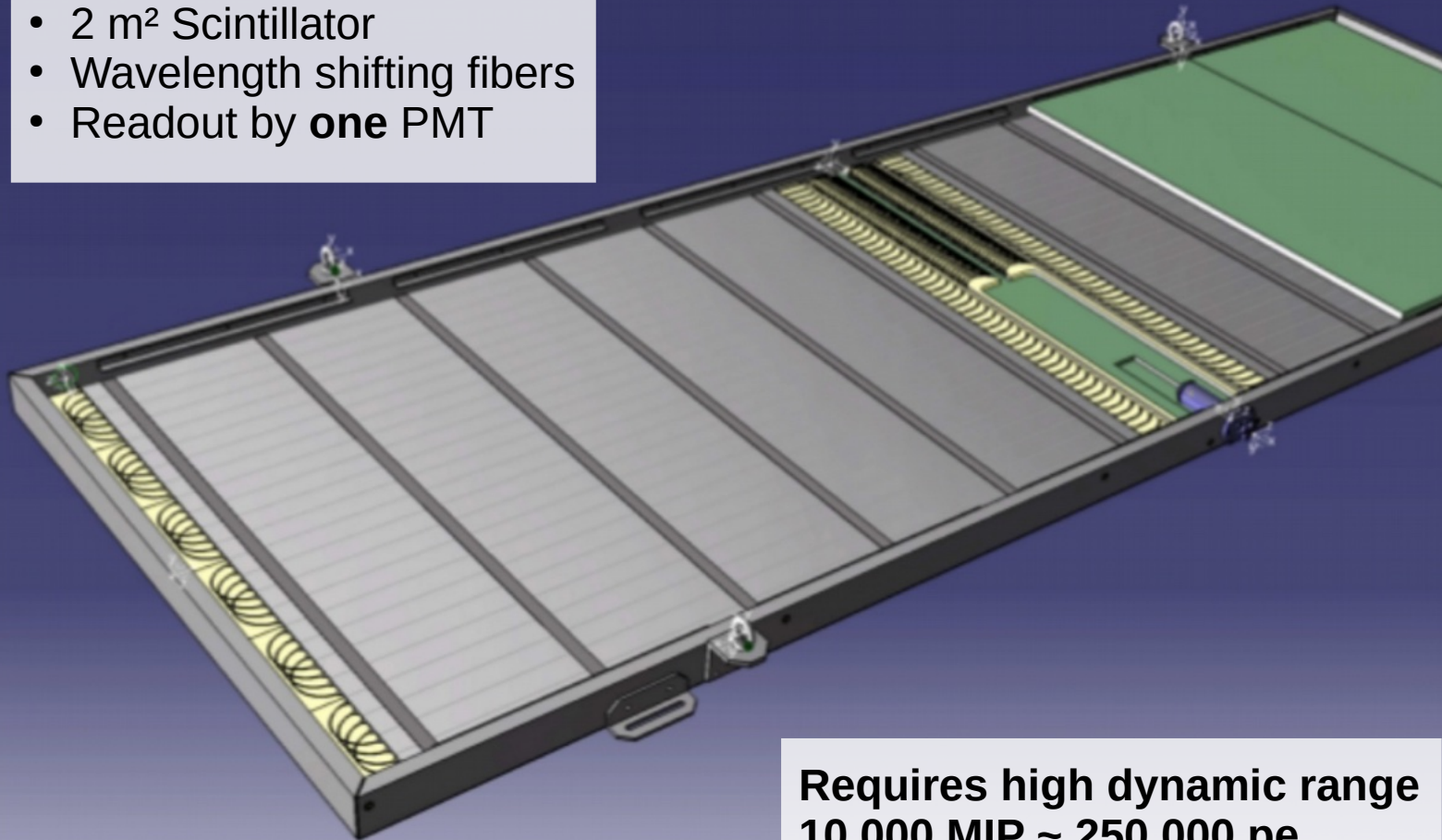
8 tiles
à 30 x 30 x 0.5 cm³



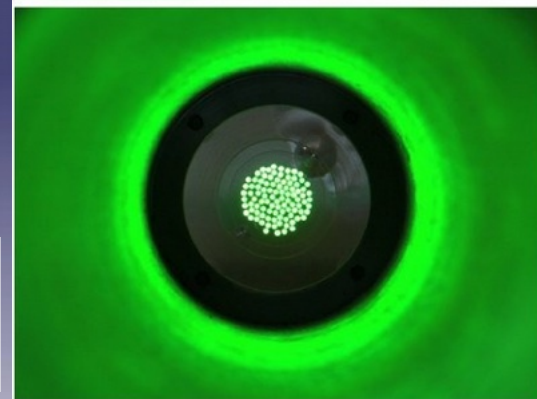
PoS ICRC2015 (2016) 596
JPS Conf. Proc. 19, 011030 (2018)

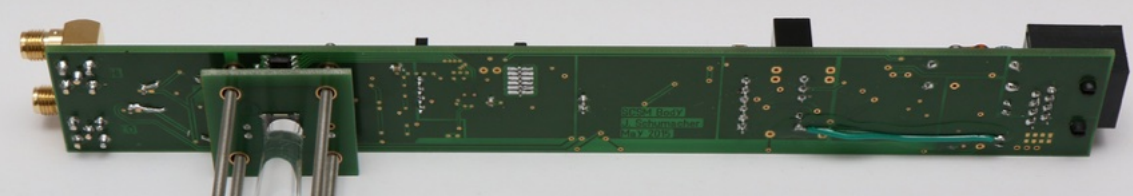
Scintillator Surface Detector (Pierre Auger)

- 2 m² Scintillator
- Wavelength shifting fibers
- Readout by **one** PMT

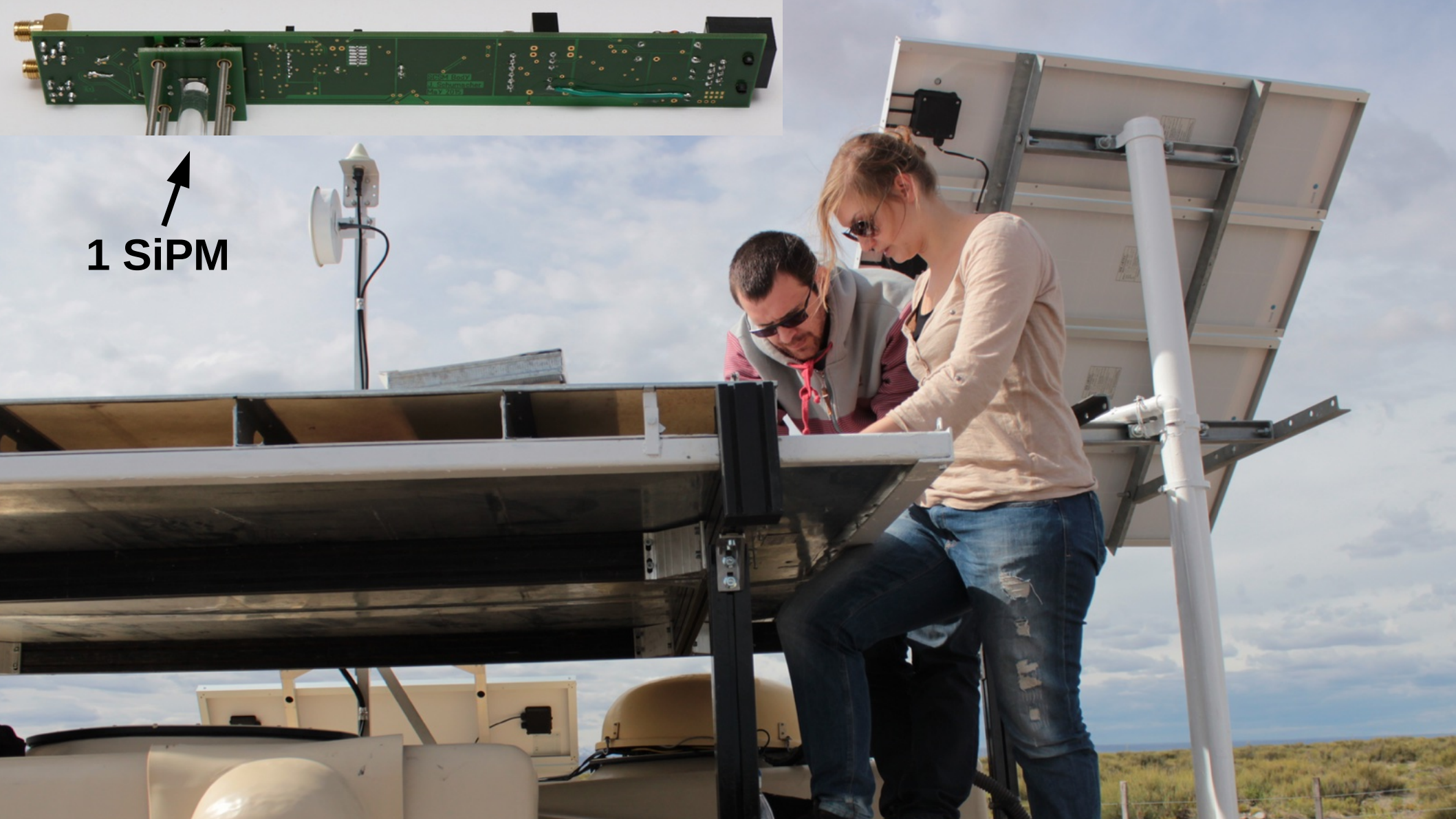


Requires high dynamic range
10,000 MIP ~ 250,000 pe



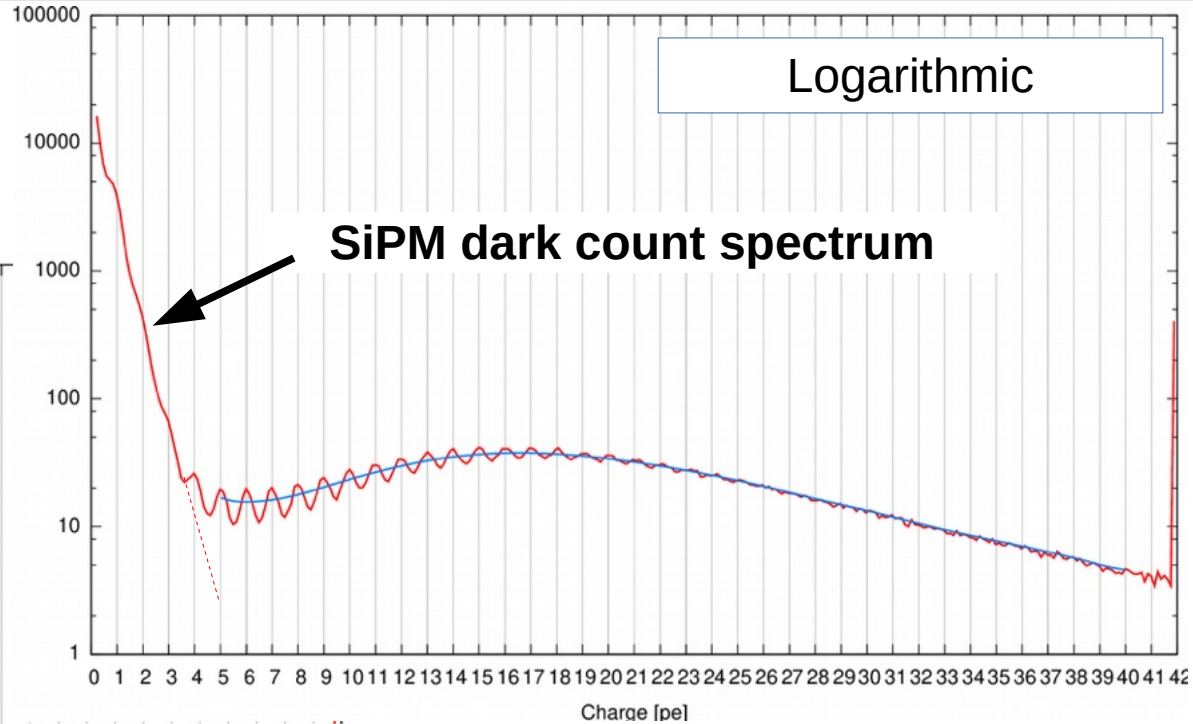
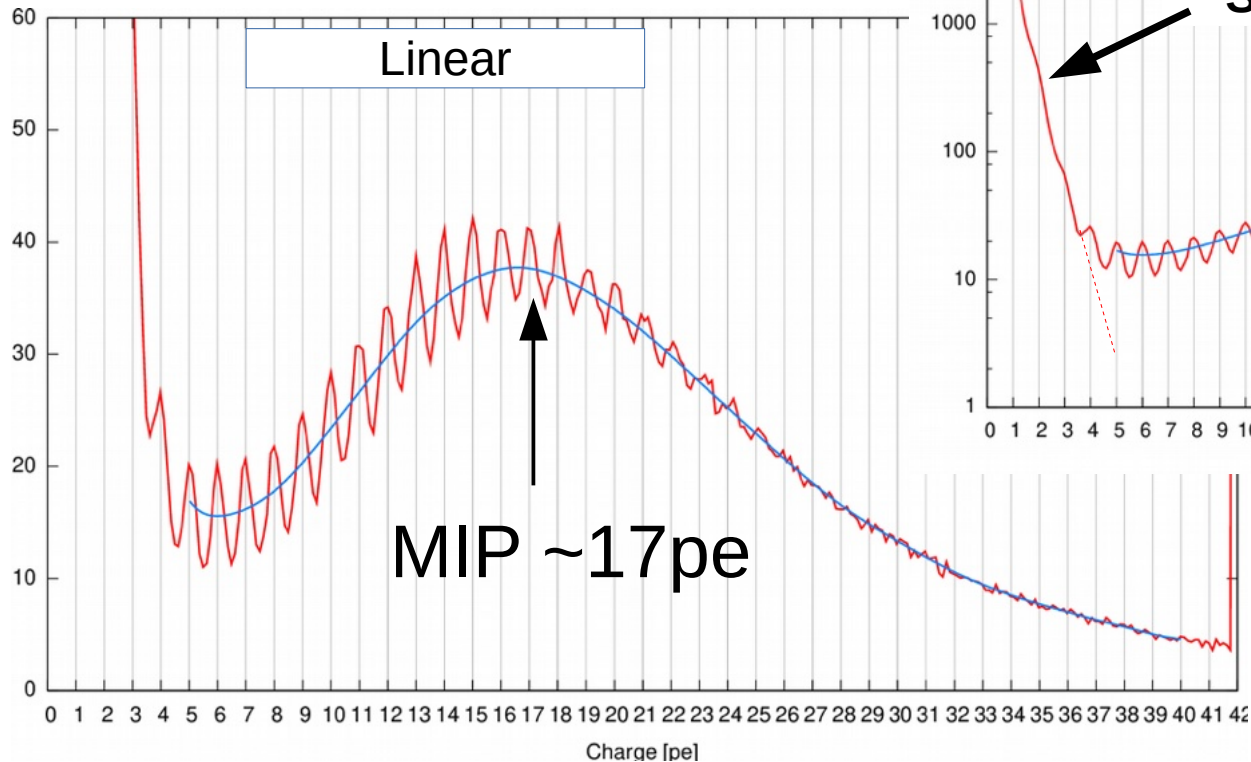


↑
1 SiPM



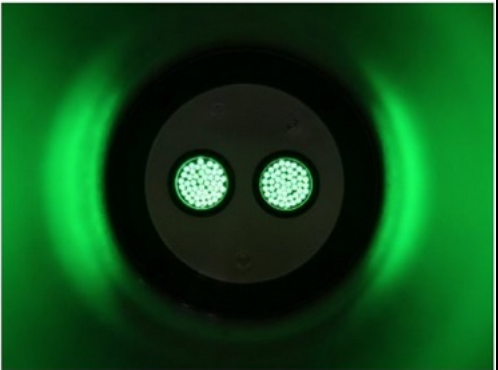
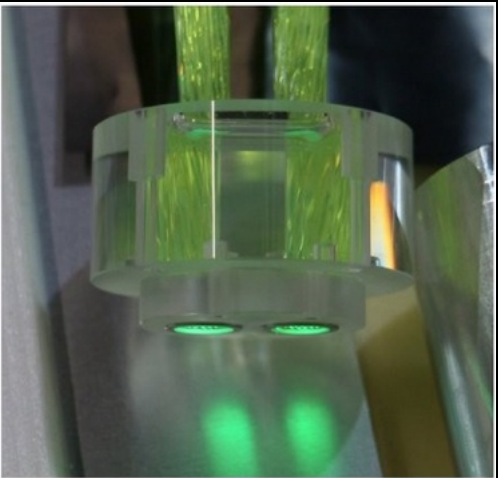
MIP Spectrum

Converted to N_{pe}
by just counting



SSD Production

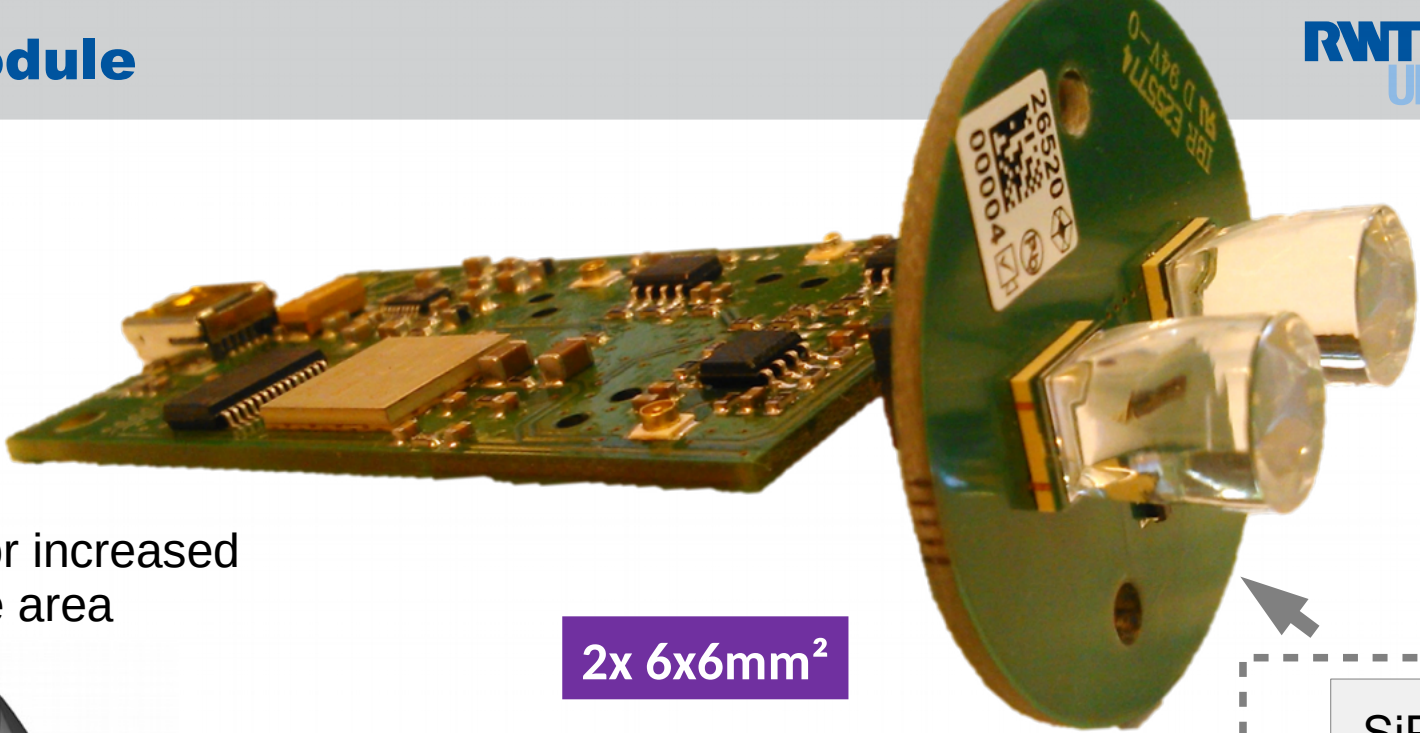
Some special detectors



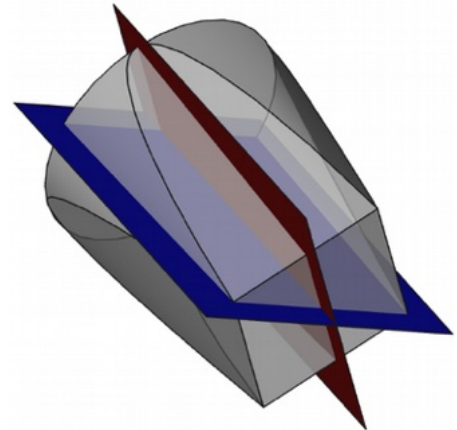
Adapt to area of SiPM



The optical module

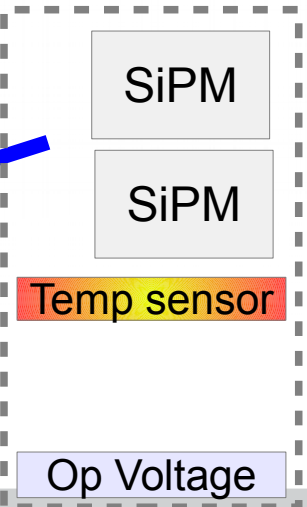
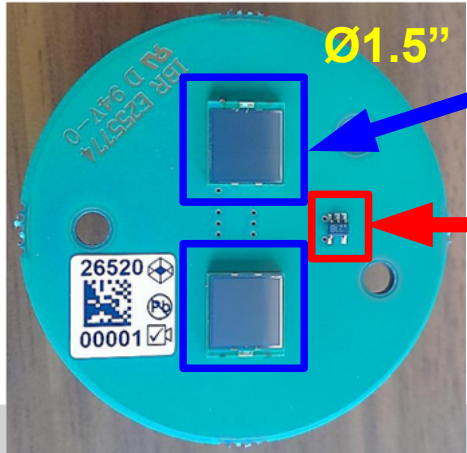


Light guides for increased sensitive area

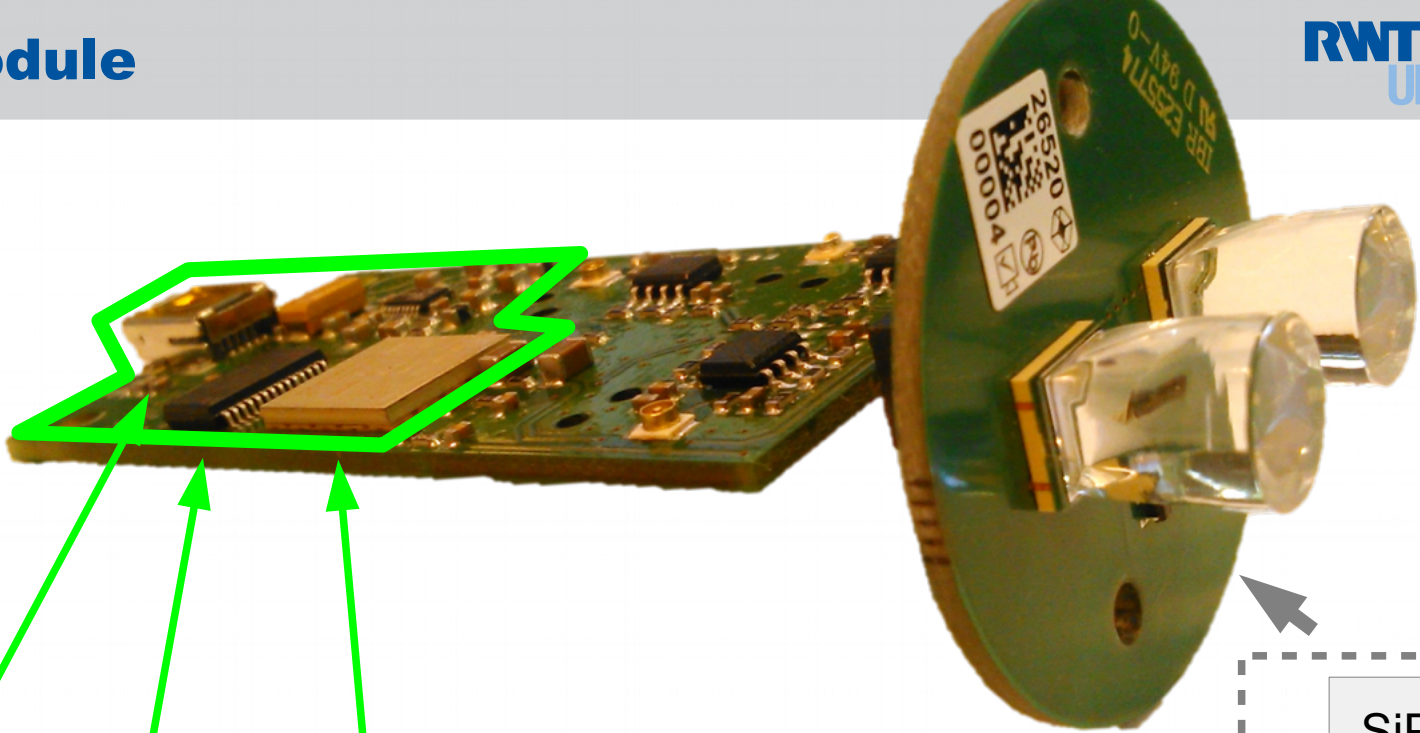


2x 6x6mm²

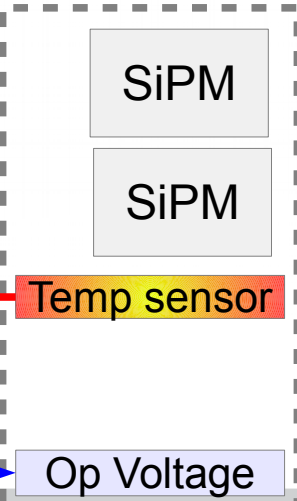
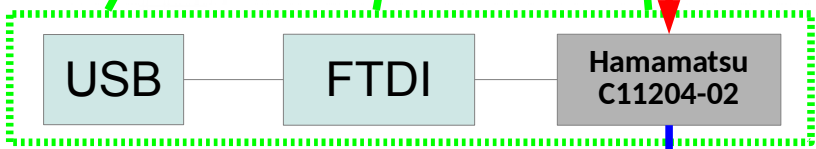
SiPMs sorted by breakdown voltage



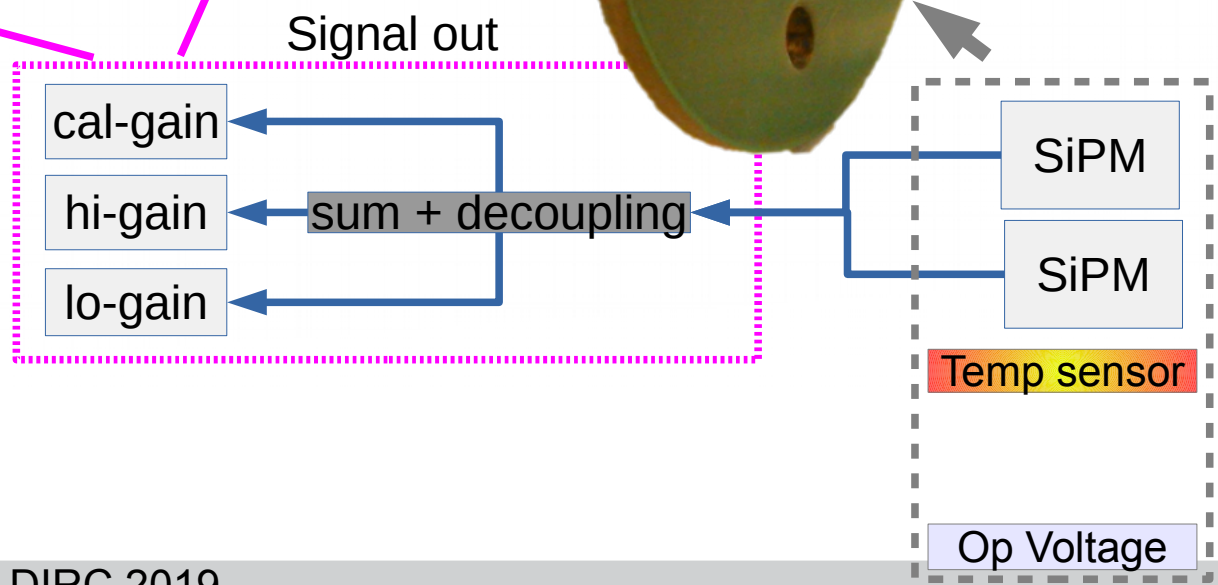
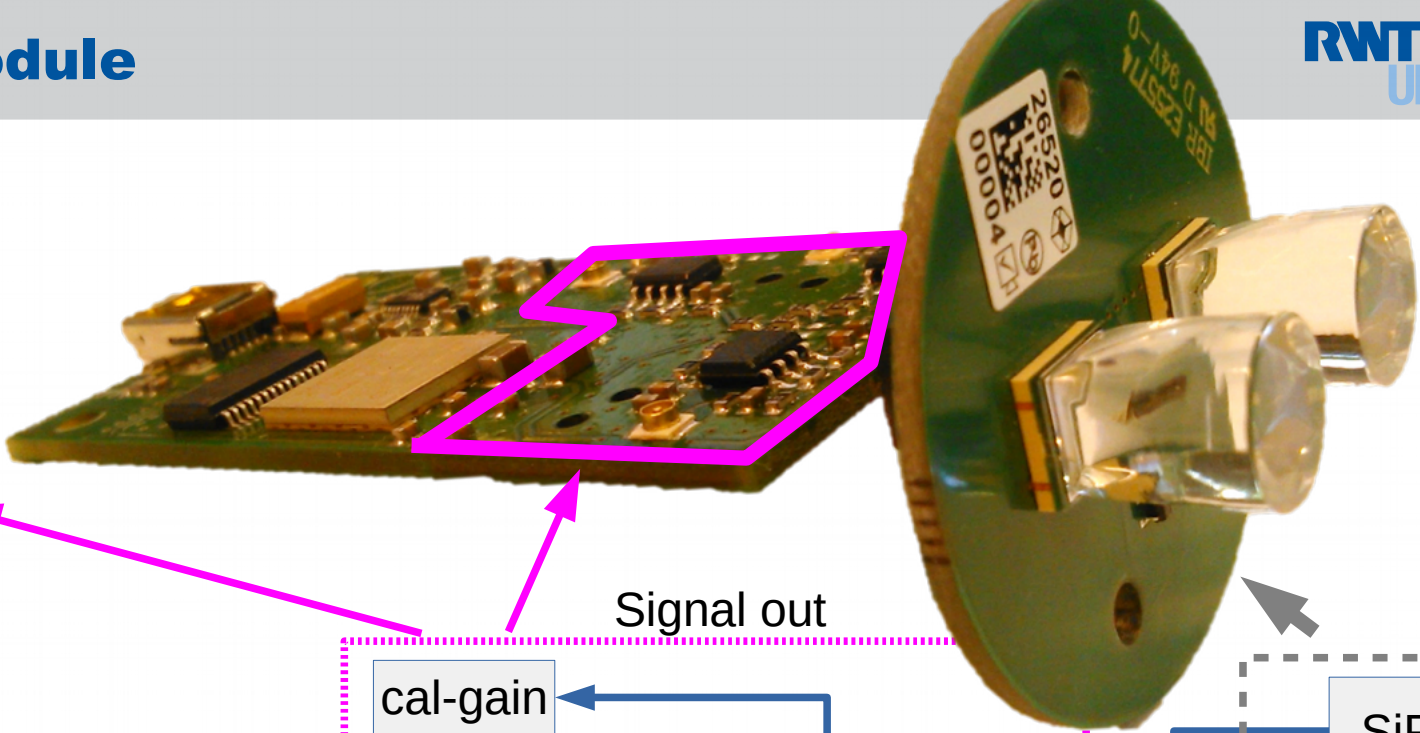
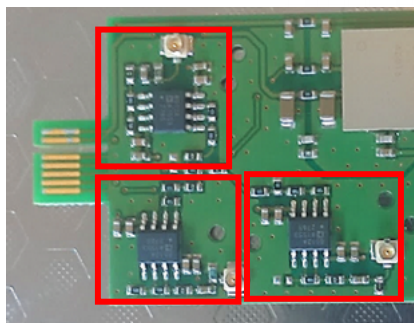
The optical module



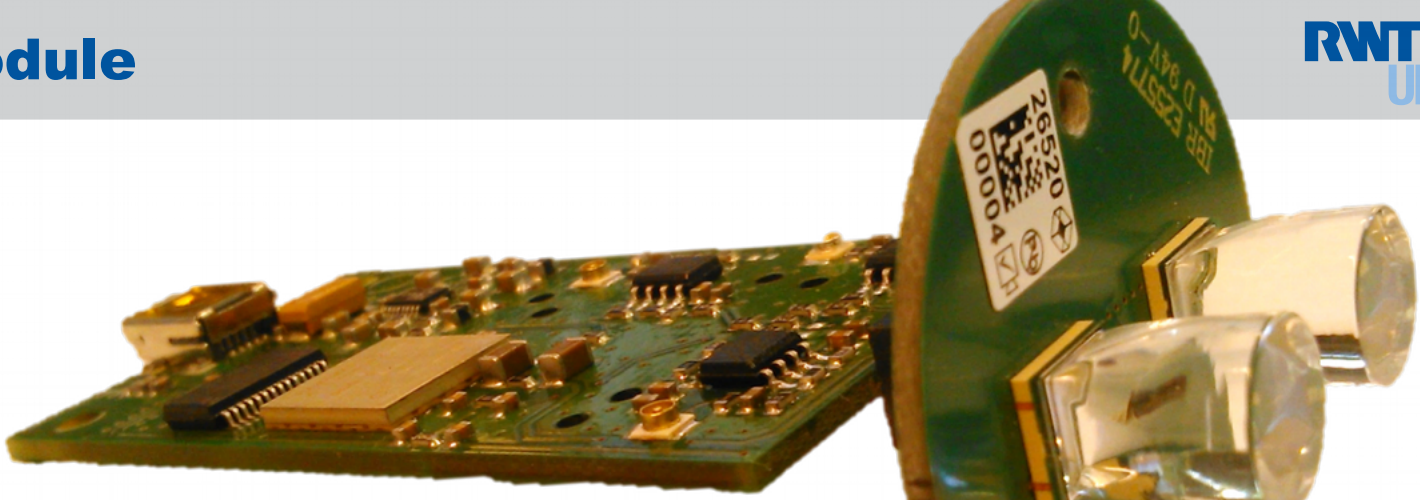
Power and communication



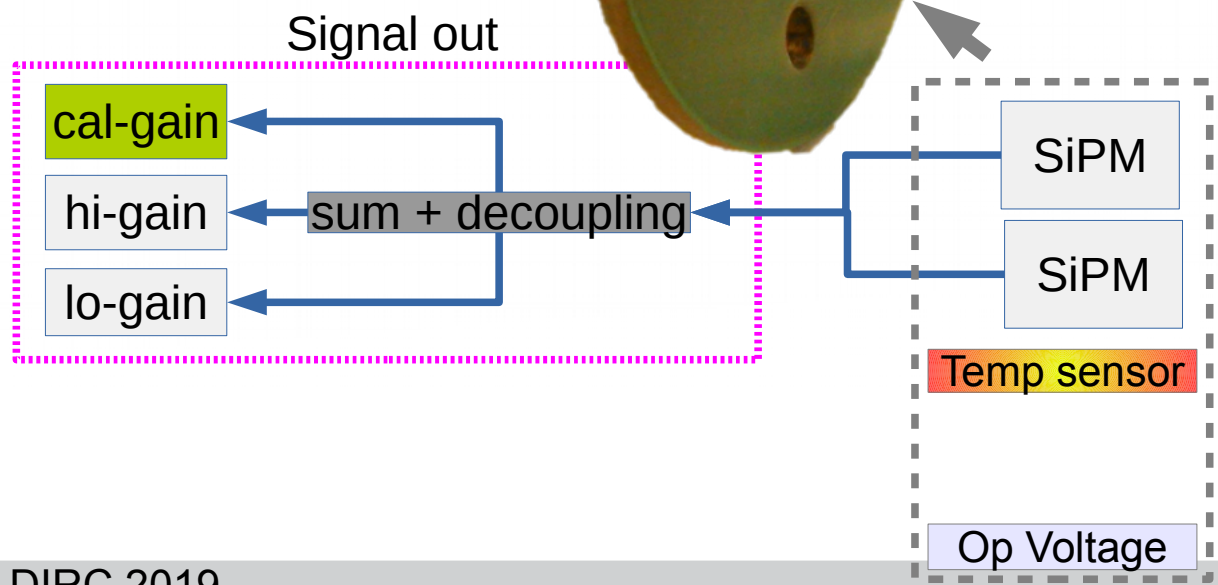
The optical module



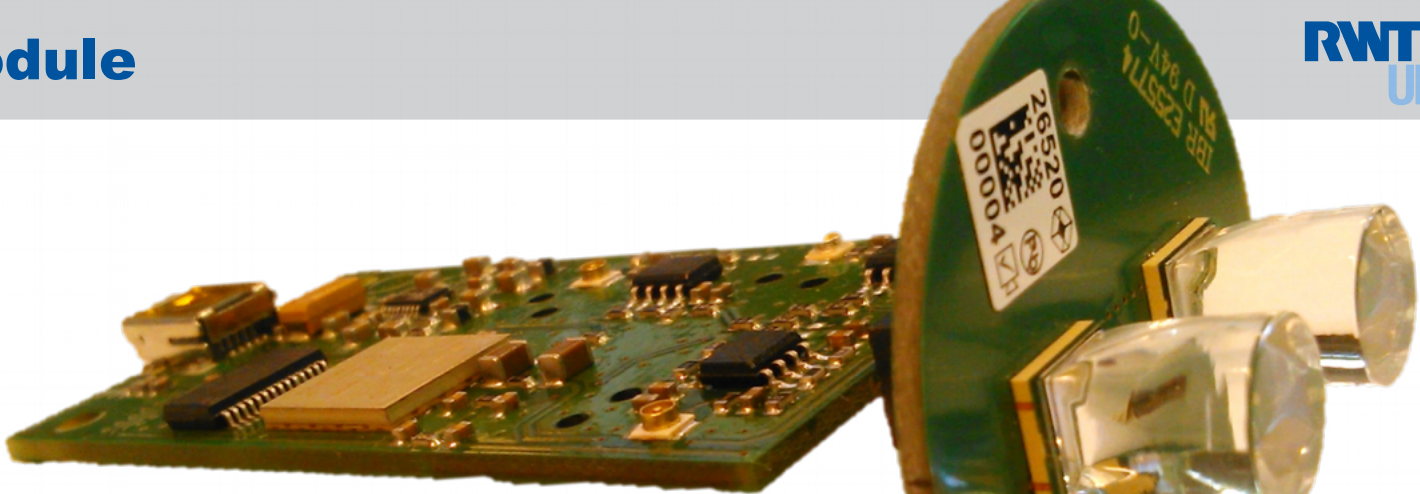
The optical module



single pe resolution

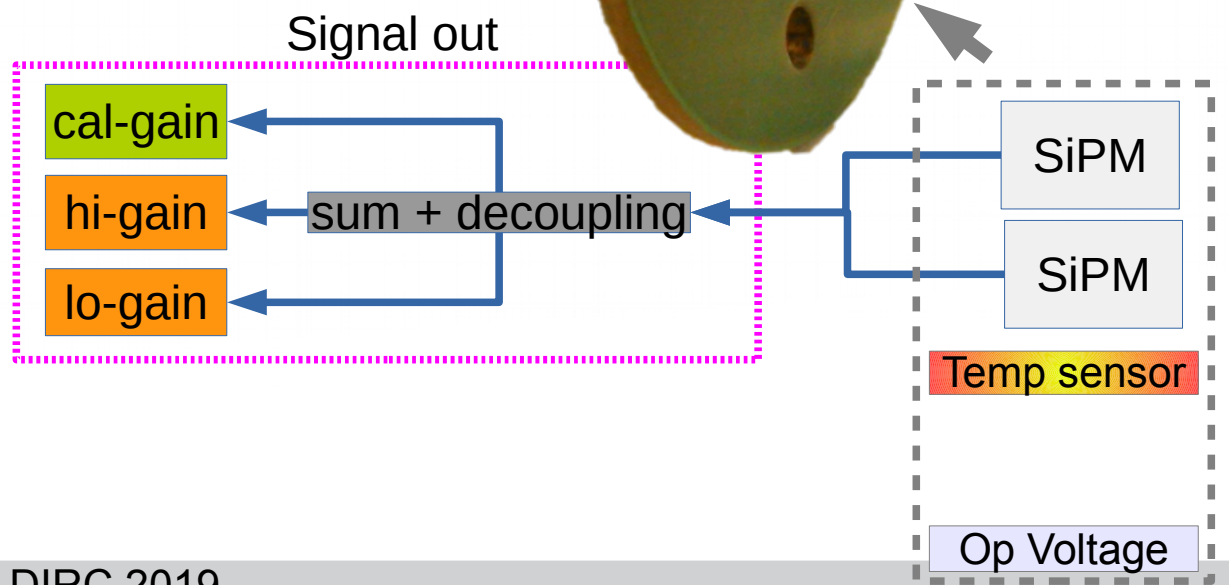


The optical module

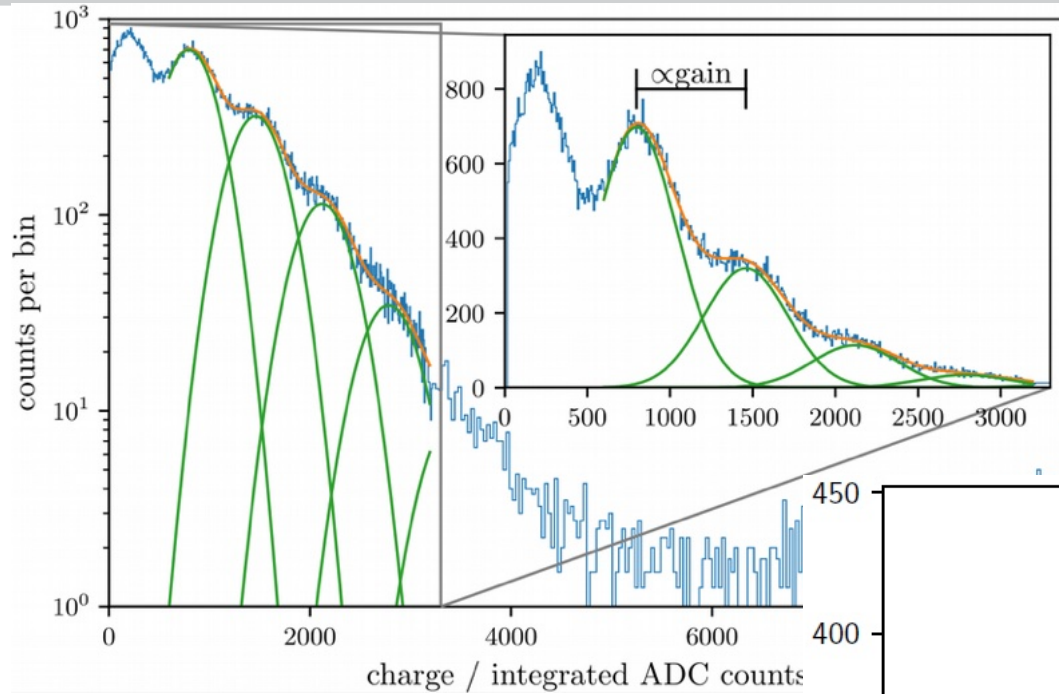


single pe resolution

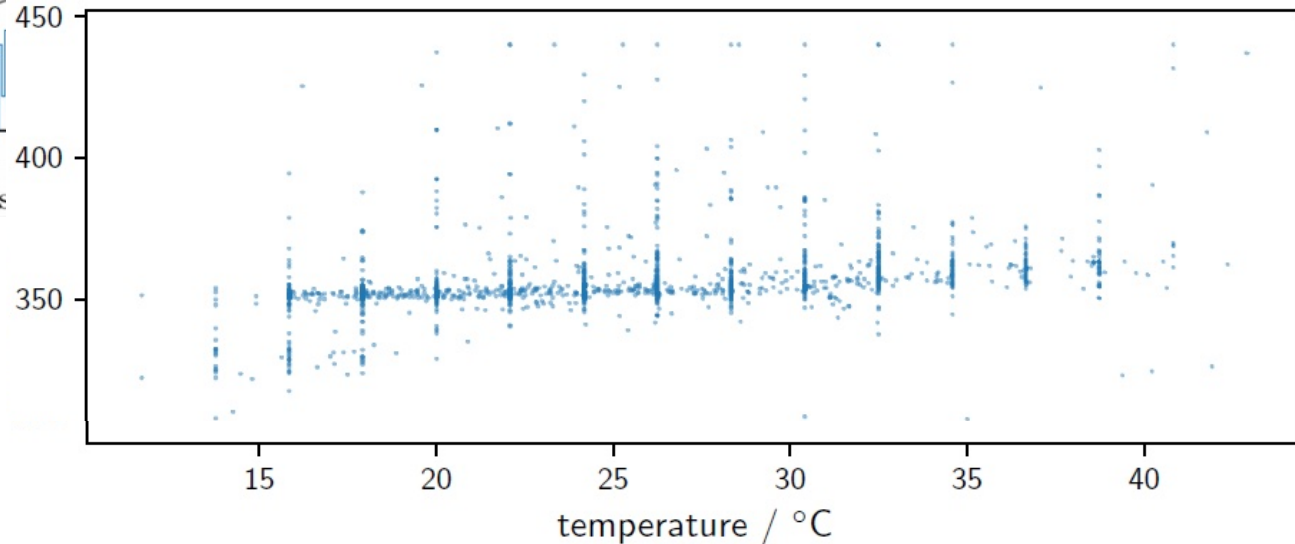
physics / MIP, high dyn. range

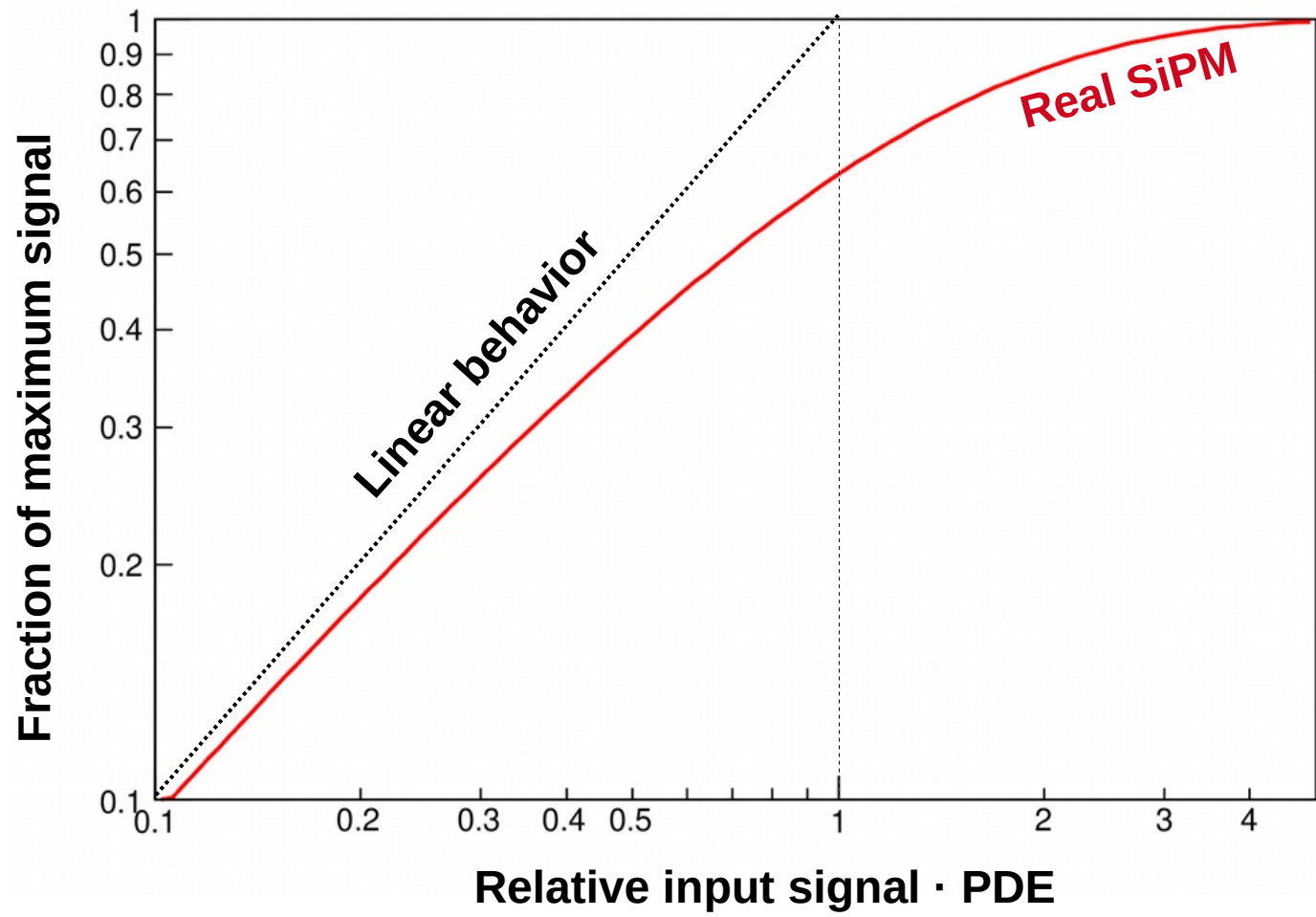


Gain in the field (Pampa Amarilla, Argentina)

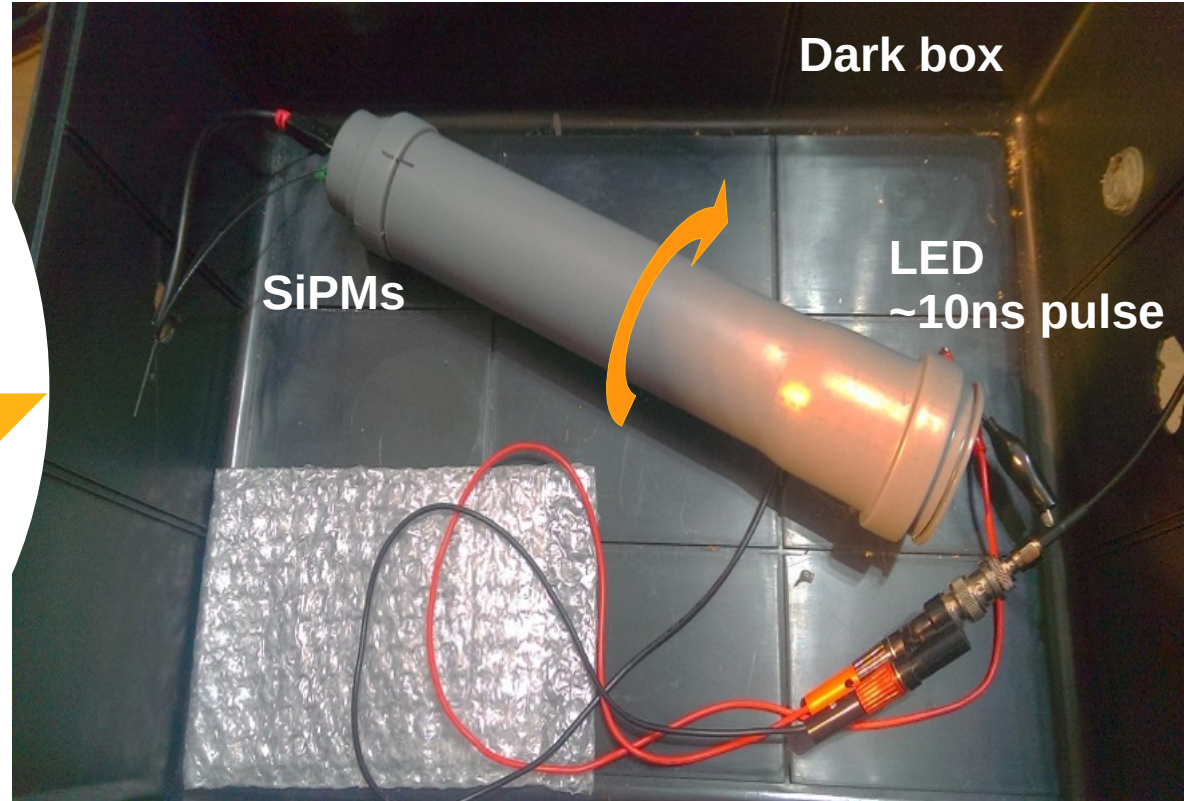
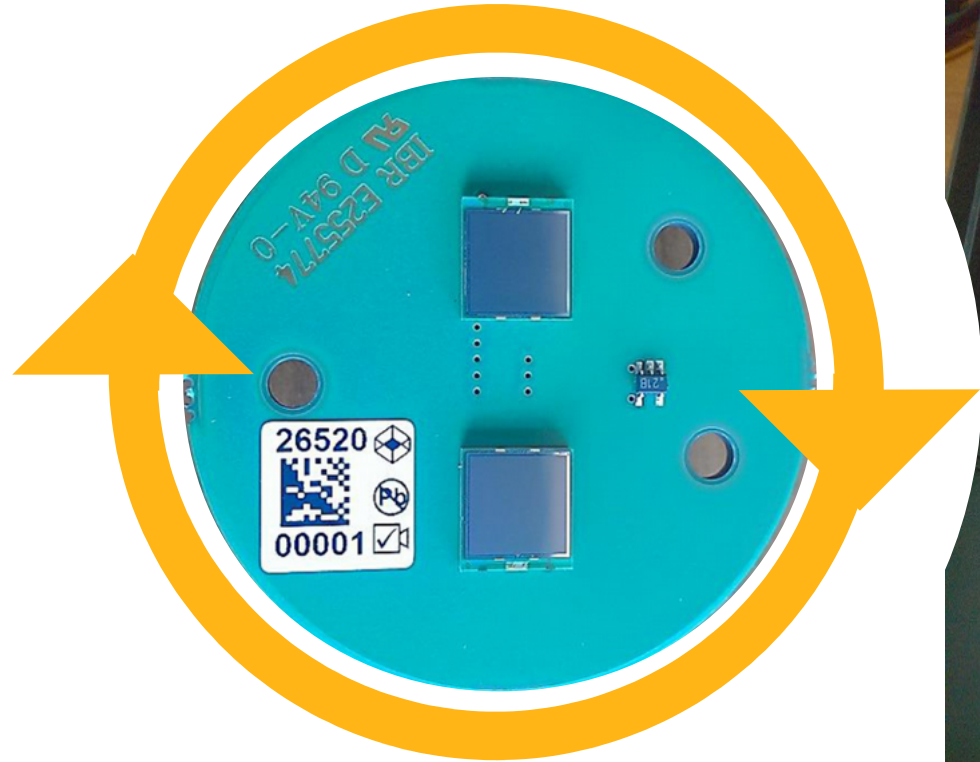


Gain vs Temperature



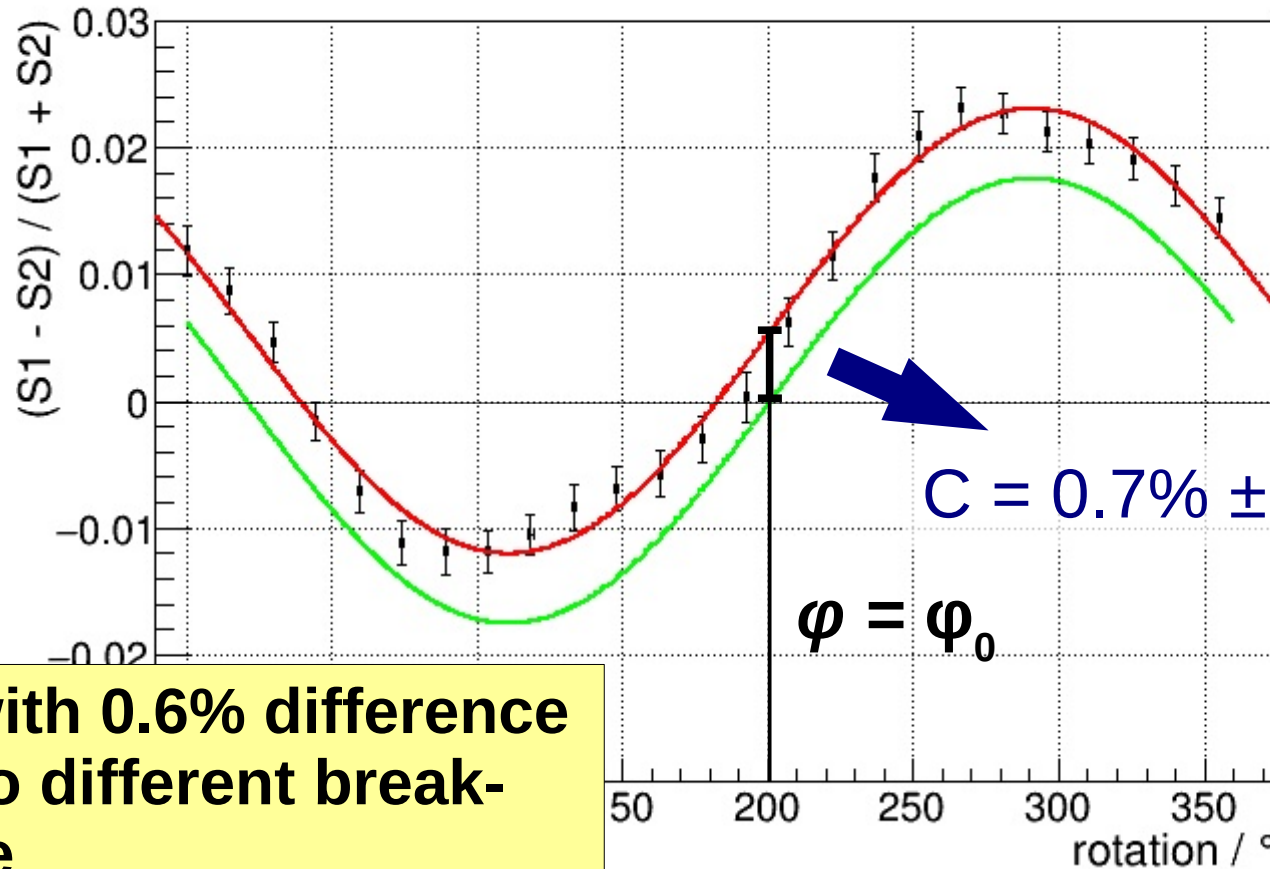


Compare two devices of same type



Determine systematic error

$$\text{Fit: } f(\varphi) = C + A \cdot \sin(\varphi - \varphi_0)$$

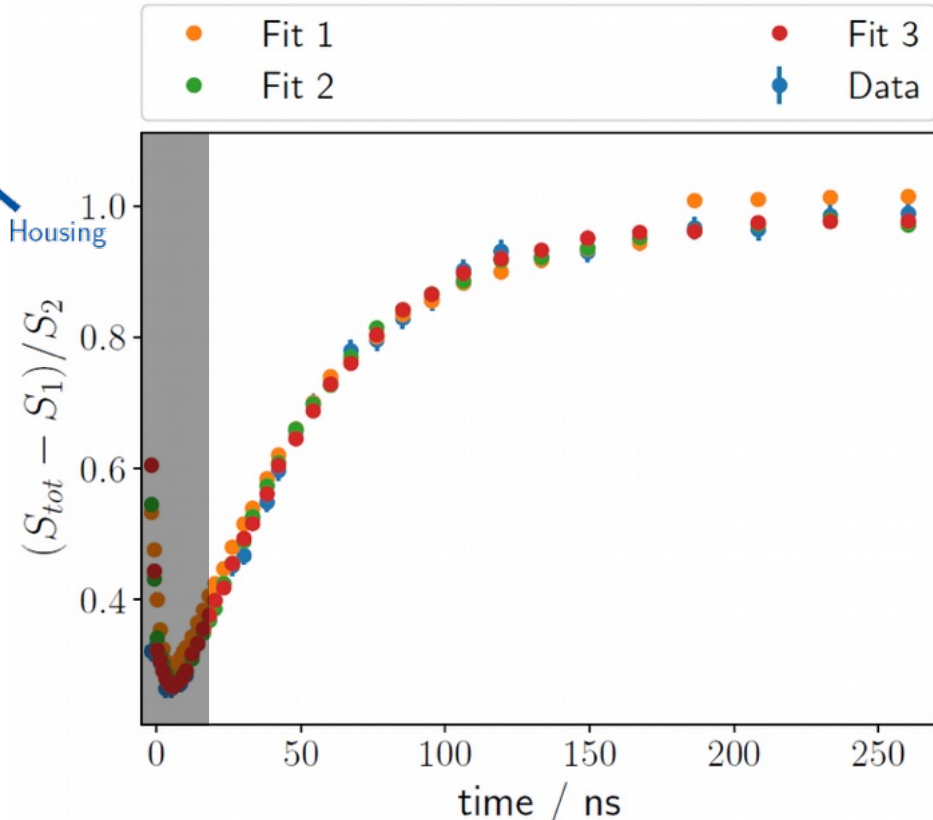
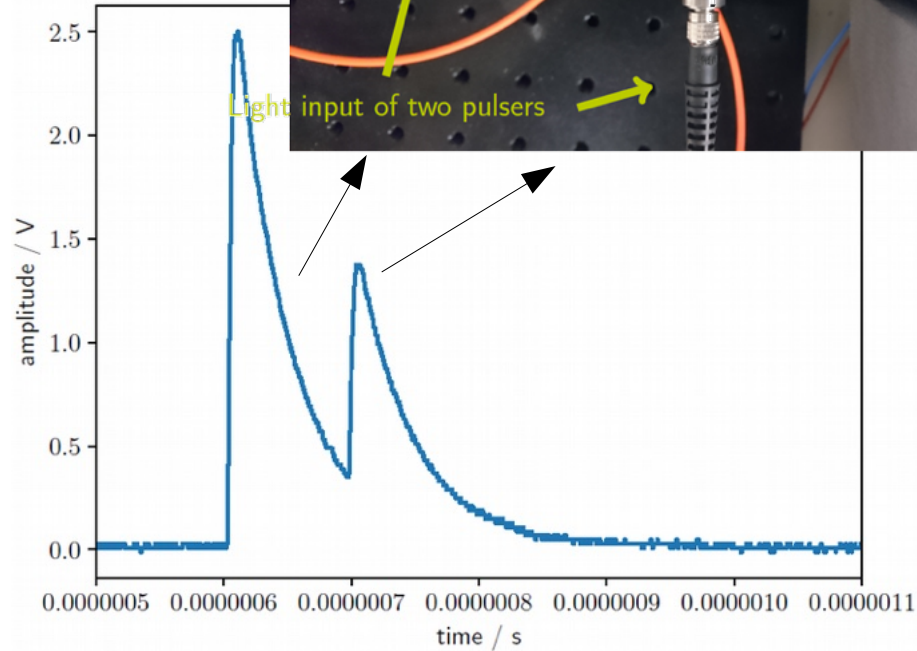
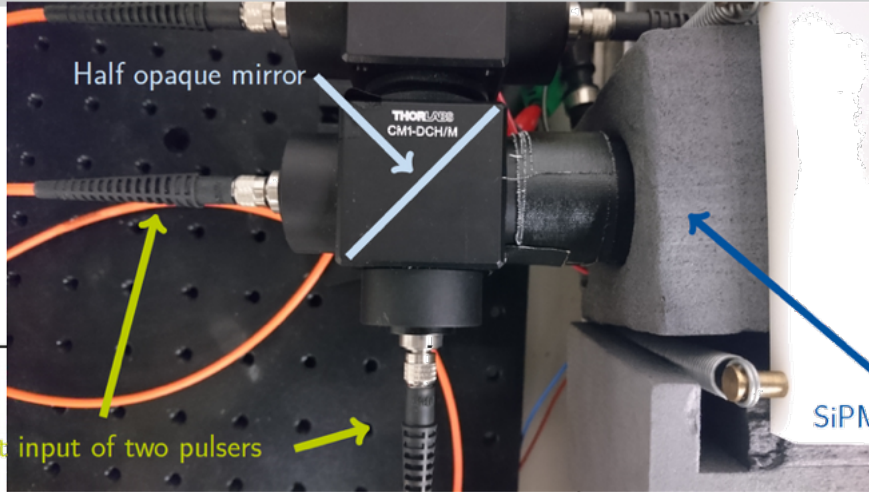


$$C = 0.7\% \pm 0.4\% \text{ (syst.)}$$

$$\varphi = \varphi_0$$

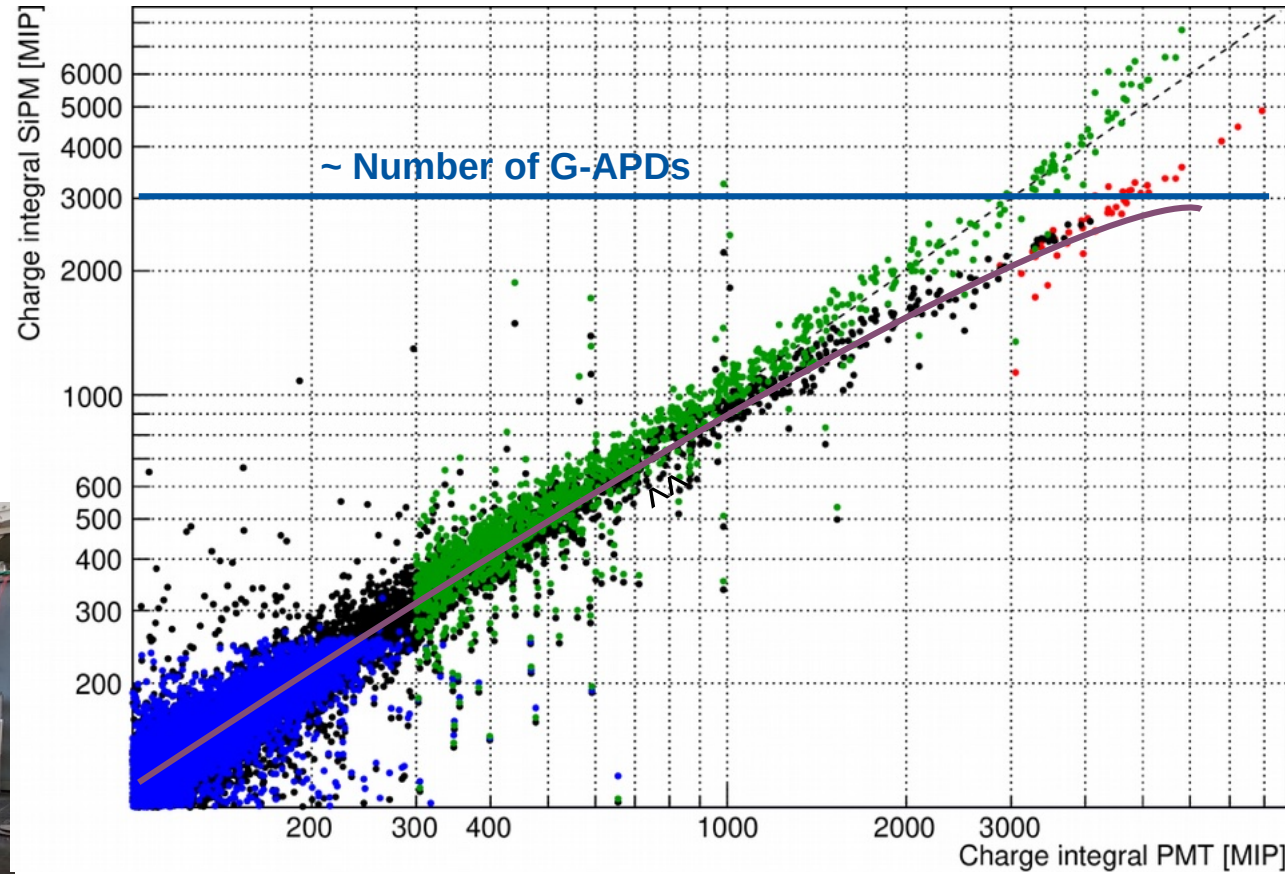
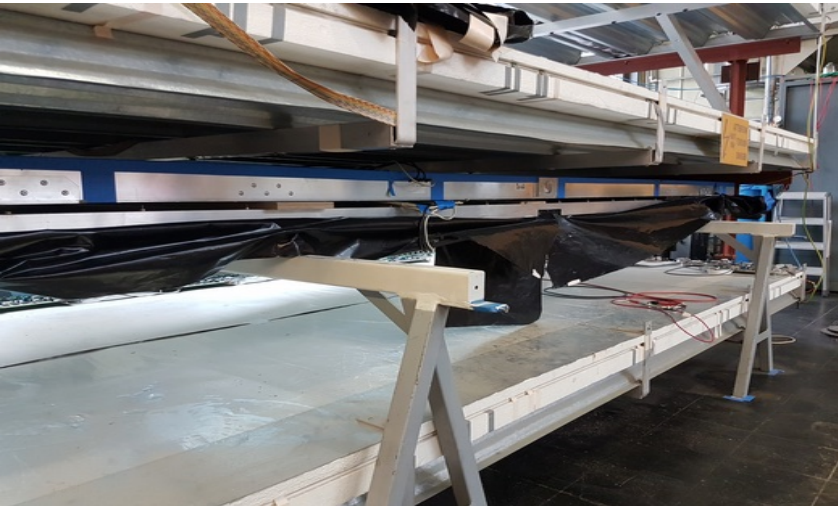
Consistent with 0.6% difference
in gain due to different break-
down voltage

Understanding the response of consecutive pulses



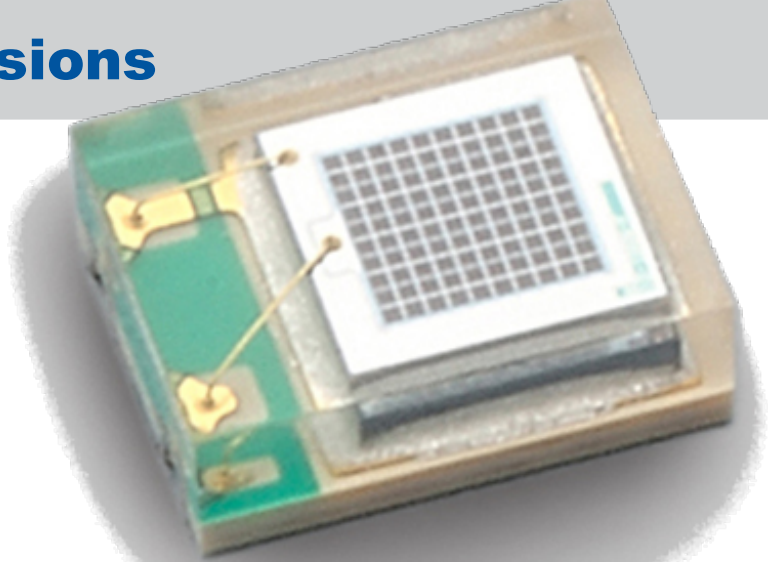
Unfolding of dynamic range

One detector with SiPM between
two detectors with PMTs

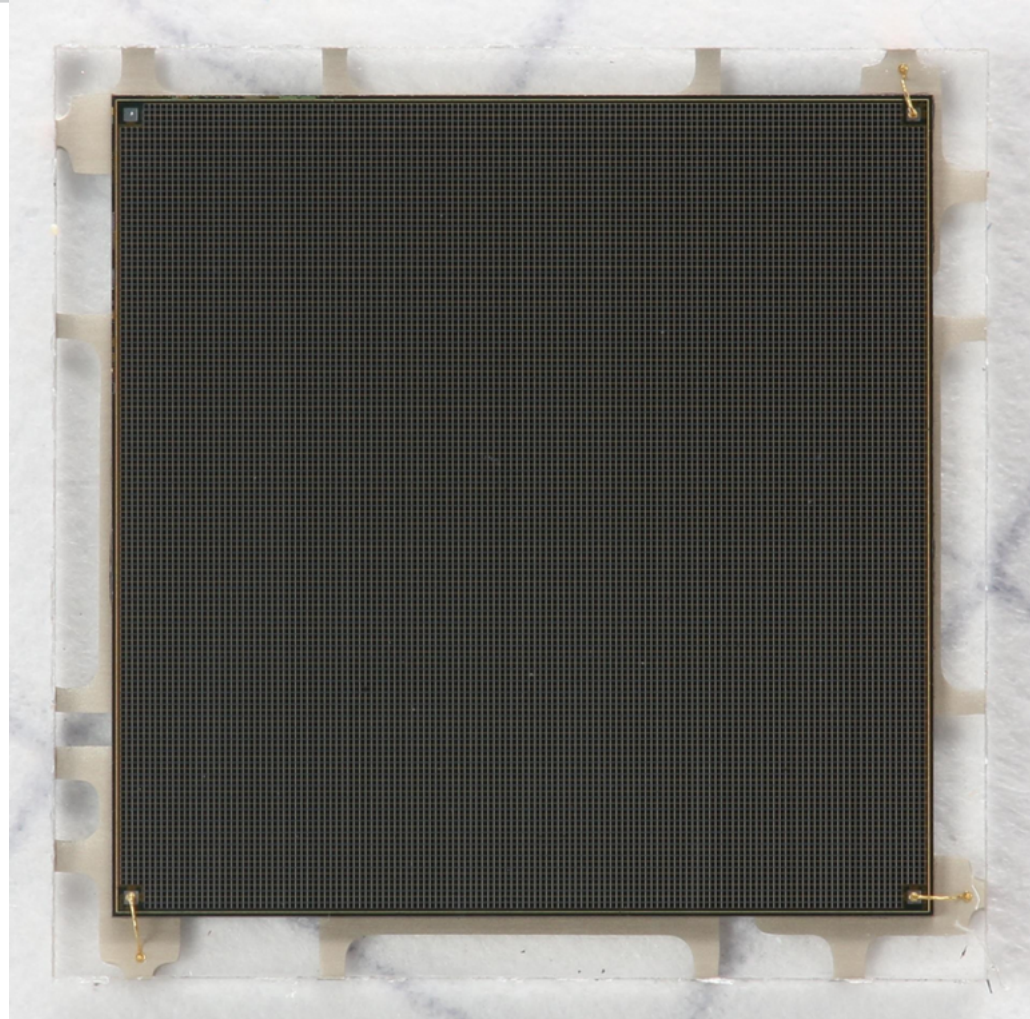


PRELIMINARY

Conclusions

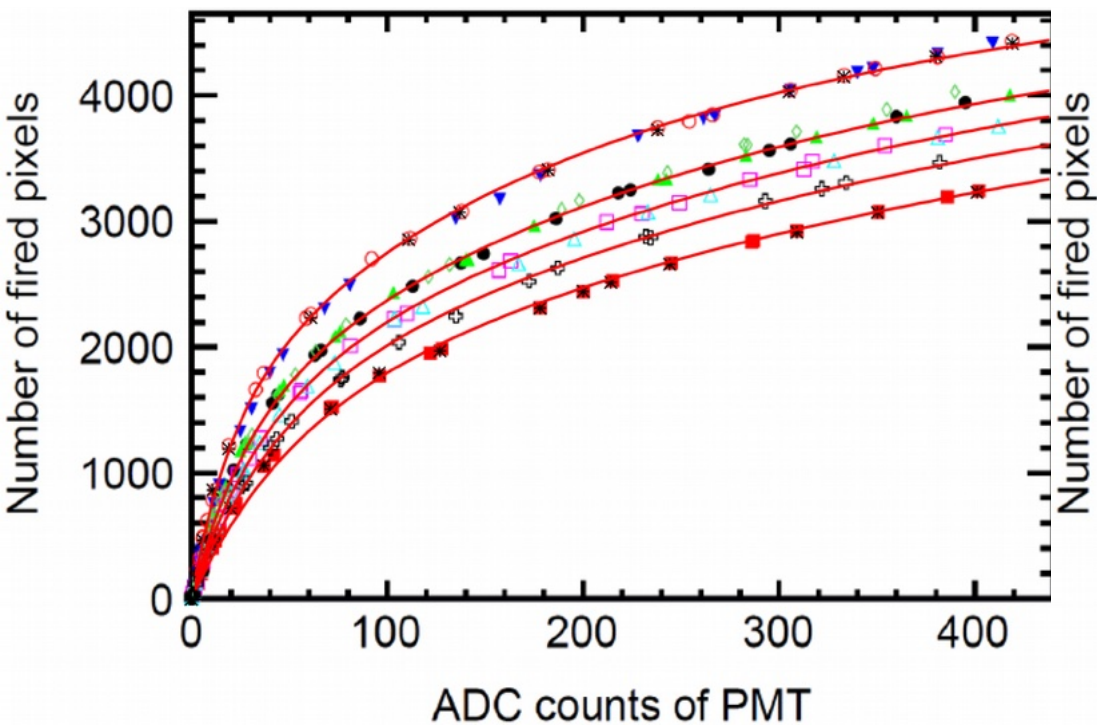


- Single photon counting
→ calibration included
- Gain as stable as precision
of power supply
- Also applicable in high
dynamic range applications



arXiv:1510.01102v4

Fit with an analytical function from paper



Simulation
(Hamamatsu 25 μ m, same parameters)

